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**Research article**

## **Classical and Bayesian inference of the weighted-exponential distribution with an application to insurance data**

**Fathy H. Riad<sup>1,2</sup>, Eslam Hussam<sup>3</sup>, Ahmed M. Gemeay<sup>4</sup>, Ramy A. Aldallal<sup>5,\*</sup>and Ahmed Z. Afify<sup>6</sup>**

<sup>1</sup> Mathematics Department, College of Science, Jouf University, P. O.Box 2014, Sakaka, Saudi Arabia

<sup>2</sup> Department of Mathematics, Faculty of Science, Minia University, Minia 61519, Egypt

<sup>3</sup> Department of Mathematics, Faculty of science Helwan university, Cairo, Egypt

<sup>4</sup> Department of Mathematics, Faculty of Science, Tanta University, Tanta 31527, Egypt

<sup>5</sup> College of Business Administration in Hotat bani Tamim, Prince Sattam Bin Abdulaziz University, Al-Kharj, Saudi Arabia

<sup>6</sup> Department of Statistics, Mathematics and Insurance, Benha University, Benha 13511, Egypt

\* Correspondence: Email: r.eldallal@psau.edu.sa.

**Abstract:** This paper addresses asymmetric flexible two-parameter exponential model called the weighted exponential (WDE<sub>x</sub>) distribution. Some of its basic mathematical features are evaluated. Its hazard rate accommodates upside-down bathtub, decreasing, decreasing-constant, increasing, and increasing-constant shapes. Five actuarial indicators are studied. We utilize nine classical and Bayesian approaches of estimation for estimating the WDE<sub>x</sub> parameters. We provide a detailed simulation study to explore and assess the asymptotic behaviors of these estimators. Two approximation methods called the Markov chain Mont Carlo and Tierney and Kadane are applied to obtain the Bayesian estimates. The efficiency and applicability of the WDE<sub>x</sub> distribution are explored by modeling a lifetime data set from insurance field, showing that the WDE<sub>x</sub> distribution provides a superior fit over its competing exponential models such as the beta-exponential, Harris extend-exponential, Marshall–Olkin exponential, Marshall–Olkin alpha-power exponential, gamma Weibull, and exponentiated-Weibull distributions.

**Keywords:** exponential distribution; risk measures; insurance data; maximum likelihood estimation; Bayes estimation; Tierney and Kadane approximation

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### **1. Introduction**

Recently, the promising attention of statisticians towards the construction of new generalized distributions has been increased to explore and develop more flexible models from the baseline models.

These generalized distributions open new horizons to study real-world problems and to provide more flexibility to model the asymmetric and complex random phenomena. Therefore, several models have been developed in the literature. One of the most handy and simplest lifetime distributions induced in the statistical literature is the exponential (Ex) distribution. It has been extensively adopted for analyzing lifetime data due to its simplicity and its lack of memory property.

On the other hand, the Ex distribution with only a constant hazard rate (HR) shape is not suitable for fitting data with different HR shapes as decreasing, unimodal (upside down bathtub), increasing or bathtub shaped hazard rates which are common in reliability and engineering, among other areas. This limitation has motivated the researchers to develop more flexible extensions of the Ex model.

For more details about other methods, the interested reader can see [1–15].

In the present article, we are motivated to develop a new flexible version of the Ex model called the weighted-exponential (WDEx) distribution, which can provide greater flexibility in fitting insurance data. Additionally, the aim of the article is three-fold: the first is devoted to studying a new form of the Ex distribution based on the weighted-G (WD-G) family [16], called the WDEx distribution. Basic distributional properties of the WDEx model are explored. The WDEx model has some desirable characteristics. It can accommodate upside-down bathtub, decreasing, decreasing-constant, increasing, and increasing-constant HR shapes. Its probability density function (PDF) possesses unimodal, right-skewed, reversed-J shape, left-skewed. The tail of the WDEx distribution is heavier than the EEx and Ex distributions as shown in Section 4 based on computational values of risk measures. The WDEx distribution can be used to model insurance data. It provides an adequate fit than other fourteen competing Ex distributions as illustrated in Section 7 for this insurance data set.

The second objective of this paper is devoted to exploring five important actuarial or risk measures, called the value at risk (VaR), tail value at risk (TVaR), tail-variance (TV), tail-variance premium (TVP), and expected shortfall (ES), for the WDEx distribution. These measures play an important role in portfolio optimization under uncertainty.

The third objective is to study the estimation of the WDEx parameters by classical and Bayesian methods. Hence, nine estimation approaches are adopted for this purpose. We explore and rank the performance of these estimators based on empirical simulation results to develop a guideline for choosing the best estimation approach for estimating the WDEx parameters.

The paper is organized in eight Sections. In Section 2, we define the WDEx model. Its basic properties are explored in Section 3. Five actuarial indicators are studies in Section 4. The model parameters are estimated using classical and Bayesian estimation in Section 5. Simulation results are provided in Section 6. Section 7 is devoted to analyzing a real-life data. Finally, some remarks are given to conclude the paper in Section 8.

## 2. The WDEx distribution

In this section, we define the two-parameter WDEx distribution based on the WD-G family of [16]. The cumulative distribution function (CDF) and PDF of the WDEx distribution has the forms

$$F(x) = \frac{\log \left[ \left( 1 - e^{-\lambda x} \right)^{\alpha} + 1 \right]}{\log(2)} \quad (2.1)$$

and

$$f(x) = \frac{\alpha \lambda e^{-\lambda x} (1 - e^{-\lambda x})^{\alpha-1}}{\log(2) [(1 - e^{-\lambda x})^\alpha + 1]}. \quad (2.2)$$

The survival function (SF), HR function (HRF) and the reversed-HRF (RHRF) are, respectively, given by

$$S(x) = 1 - \frac{\log[(1 - e^{-\lambda x})^\alpha + 1]}{\log(2)}, \quad (2.3)$$

$$h(x) = \frac{\alpha \lambda (1 - e^{-\lambda x})^\alpha}{(e^{\lambda x} - 1) [(1 - e^{-\lambda x})^\alpha + 1] \{\log(2) - \log[(1 - e^{-\lambda x})^\alpha + 1]\}} \quad (2.4)$$

and

$$r(x) = \frac{\alpha \lambda e^{-\lambda x} (1 - e^{-\lambda x})^{\alpha-1}}{((1 - e^{-\lambda x})^\alpha + 1) \log((1 - e^{-\lambda x})^\alpha + 1)}. \quad (2.5)$$

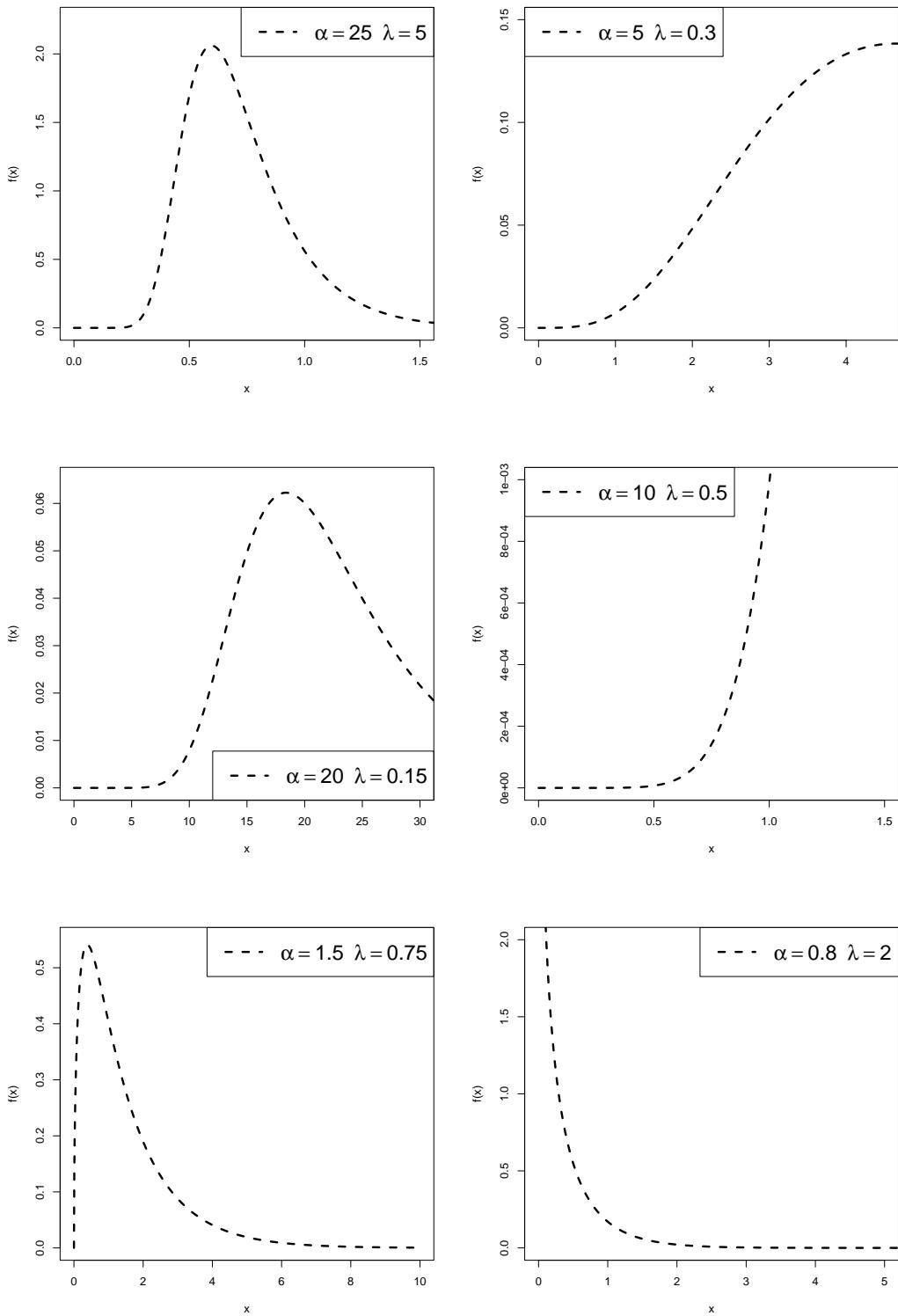
The limiting behaviors of the PDF (2.2) and HRF (2.4) are given by

$$\lim_{x \rightarrow 0} f(x) = \begin{cases} \infty, & \alpha < 1 \\ 0, & \alpha \geq 1 \end{cases}, \quad \lim_{x \rightarrow \infty} f(x) = 0.$$

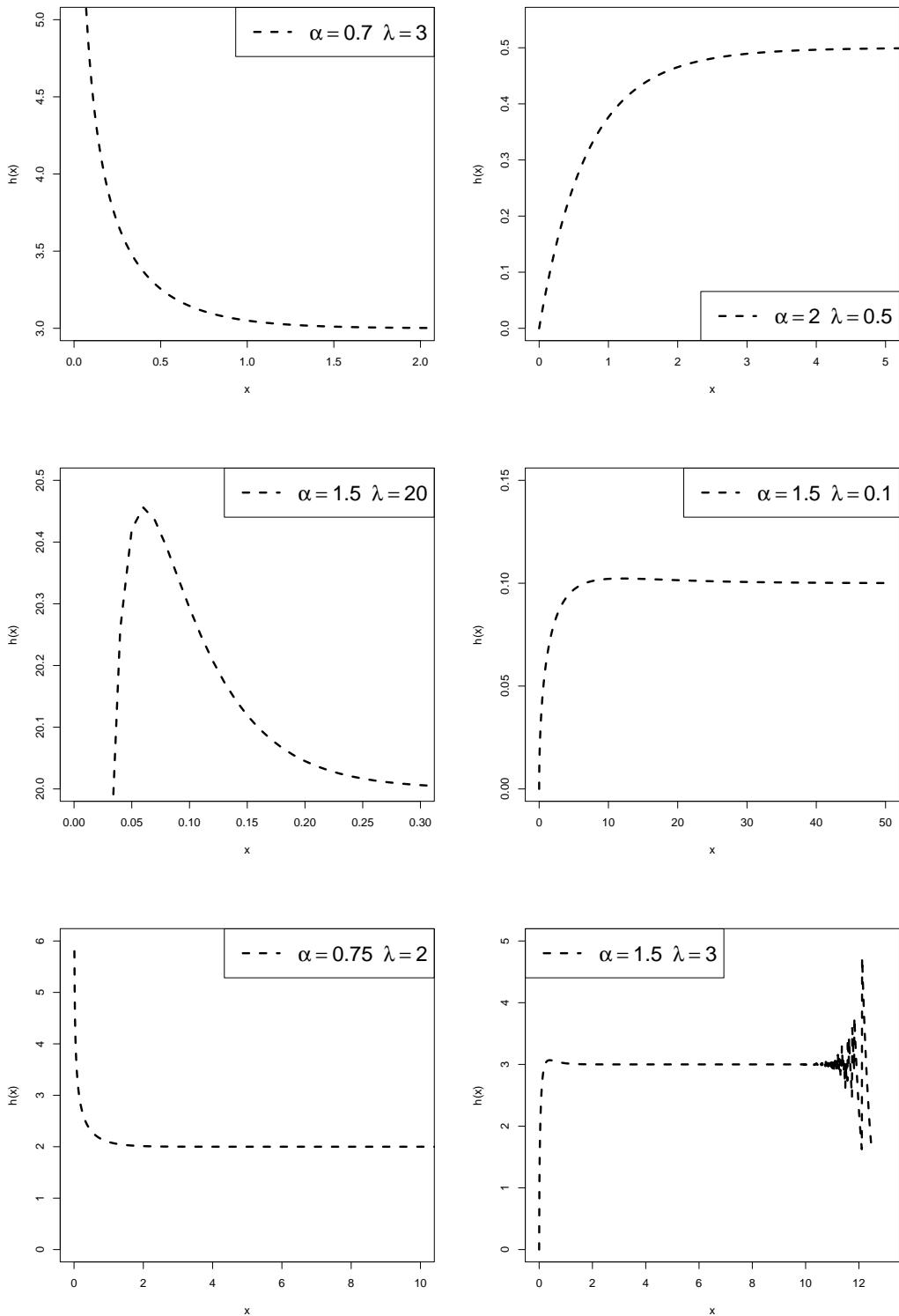
and

$$\lim_{x \rightarrow 0} h(x) = \begin{cases} \infty, & \alpha < 1 \\ 0, & \alpha \geq 1 \end{cases}, \quad \lim_{x \rightarrow \infty} h(x) = \lambda.$$

Plots of the PDF and HRF of the WDEEx distribution are illustrated in Figures 1 and 2, respectively. These graphs show that the WDEEx model possesses decreasing-constant, upside-down bathtub, increasing, decreasing, and increasing-constant HR functions along with sudden spikes at the end of upside-down bathtub shape, which manages the model to analyze a non-stationary real-life data.



**Figure 1.** Plots of the PDF of the WDEx distribution with various parametric values.

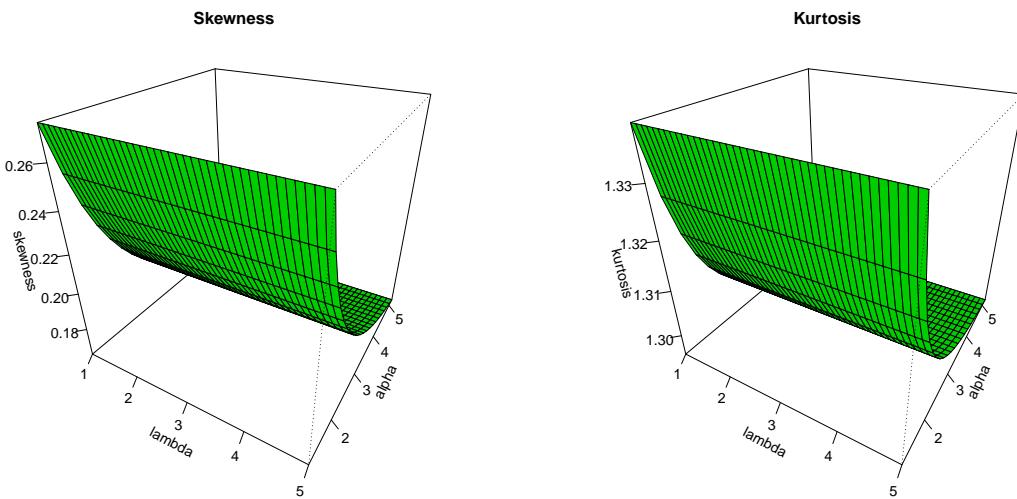


**Figure 2.** Plots of the HRF of the WDEx distribution with various parametric values.

The quantile function (QF) of the WDEEx distribution is given by

$$Q(u) = -\frac{\log[1 - (2^u - 1)^{1/\alpha}]}{\lambda}, \quad (2.6)$$

we can determine first, second and third quartile of the WDEEx distribution by setting  $u = 0.25, 0.5, 0.75$ , respectively. Also, by using Eq (2.6), we can determine Bowley's skewness and Moor's kurtosis for the WDEEx distribution. The WDEEx skewness and kurtosis graphs for the different parametric values of  $\alpha$  and  $\lambda$  are shown in Figure 3.



**Figure 3.** Plots of the skewness and kurtosis of the WDEEx model.

### 3. Probabilistic and statistical properties

In this section, we provide some mathematical properties of the WDEEx distribution.

#### 3.1. Linear representation

A useful linear representations of the CDF and PDF of the WDEEx model are introduced in this subsection. For  $-1 < x \leq 1$ , we have

$$\log(1 + x) = \sum_{k=1}^{\infty} (-1)^{k+1} \frac{x^k}{k}. \quad (3.1)$$

By applying the power series (3.1) to (2.1), the CDF and PDF of the WDEEx distribution can be expressed as

$$F(x; \alpha, \lambda) = \sum_{k=1}^{\infty} \Phi_k G_k(x)$$

and

$$f(x; \alpha, \lambda) = \sum_{k=1}^{\infty} \Phi_k g_k(x),$$

where  $\Phi_k = \frac{(-1)^{k+1}}{k \log(2)}$ ,  $G_k(x) = (1 - e^{-\lambda x})^{\alpha k}$  and  $g_k(x) = \alpha k \lambda e^{-\lambda x} (1 - e^{-\lambda x})^{\alpha k - 1}$  are the CDF and PDF of the exponentiated-exponential (EEx) distribution with parameters  $\lambda$  and  $\alpha k$ .

### 3.2. Moments and generating function

The  $r$ th moments of the WDEEx distribution takes the form

$$\mu'_r = E(X^r) = \int_0^\infty x^r f(x) dx = \sum_{k=1}^{\infty} \sum_{i=0}^{\infty} \Phi_k \frac{\alpha k \Gamma(r+1)}{\lambda^r} (-1)^i \frac{D(\alpha k - 1, i)}{(i+1)^{k+1}},$$

where  $D(a-1, b) = \frac{(a-1)\dots(a-b)}{b!}$ . Replacing  $r = 1, 2, 3$ , and  $4$  in the last formula, we obtain the first four moments for the WDEEx distribution.

The  $n$ th moment residual life of the WDEEx distribution,  $E(X^n|X > x)$ , has the form

$$\begin{aligned} E(X^n|X > x) &= \int_x^\infty y^n f(y) dy = \sum_{k=1}^{\infty} \Phi_k \frac{\alpha \lambda k}{1 - (1 - e^{-\lambda x})^{\alpha k}} \int_x^\infty y^n \exp(-\lambda y) (1 - \exp(-\lambda y))^{\alpha k - 1} dy \\ &= \sum_{k=1}^{\infty} \Phi_k \frac{\alpha (-1)^n}{\lambda^n [1 - (1 - e^{-\lambda x})^{\alpha k}]} \int_0^{\exp(-\lambda x)} (\log z)^n (1 - z)^{\alpha k - 1} dz, \end{aligned}$$

where  $z = \exp(-\lambda y)$  and the mean residual life function of the proposed model is  $E(X|X > x) - 1$ .

the moment generating function (MGF) of the WDEEx distribution is

$$M(t) = \int_0^\infty e^{tx} f(x) dx = \sum_{k=1}^{\infty} \Phi_k \alpha k \beta \left(1 - \frac{t}{\lambda}, \alpha k\right),$$

where  $\beta(a, b)$  is beta function. The characteristic function of the WDEEx model is

$$\phi(t) = \int_0^\infty e^{itx} f(x) dx = \sum_{k=1}^{\infty} \Phi_k \alpha k \beta \left(1 - \frac{it}{\lambda}, \alpha k\right).$$

### 3.3. Mean deviation, bonferroni and Lorenz curves

The mean deviation about the mean of the WDEEx model reduces to

$$\begin{aligned} \psi_1 &= \int_0^\infty |x - \mu| f(x) dx = 2\mu F(\mu) - 2\mu + 2 \int_\mu^\infty x f(x) dx \\ &= 2\mu F(\mu) - 2\mu + 2 \sum_{k=1}^{\infty} \Phi_k \int_\mu^\infty x g_k(x) dx = 2\mu F(\mu) - 2\mu \end{aligned}$$

$$+ \frac{2\alpha k}{\lambda} \sum_{k=1}^{\infty} \sum_{i=0}^{\infty} \Phi_k \frac{(-1)^i}{(i+1)^2} \binom{\alpha k - 1}{i} [1 + \mu \lambda (i+1)] \exp(-\mu \lambda (i+1)).$$

The mean deviation about the median of the WDEx distribution takes the form

$$\begin{aligned} \psi_2 &= \int_0^{\infty} |x - M| f(x) dx = 2 \int_M^{\infty} x f(x) dx - \mu = 2 \sum_{k=1}^{\infty} \Phi_k \int_M^{\infty} x g_k(x) dx - \mu \\ &= \frac{2\alpha k}{\lambda} \sum_{k=1}^{\infty} \sum_{i=0}^{\infty} \Phi_k \frac{(-1)^i}{(i+1)^2} \binom{\alpha k - 1}{i} [1 + M \lambda (i+1)] \exp(-M \lambda (i+1)) - \mu. \end{aligned}$$

The Bonferroni and Lorenz curves for a positive random variable  $X$  are defined the following formulae

$$\begin{aligned} B(F(x)) &= \frac{E(X|X \leq x)}{\mu} \\ \text{and } L(F(x)) &= \frac{E(X|X \leq x)F(x)}{\mu}, \end{aligned}$$

where the quantity  $E(X|X \leq x)$  has the form

$$E(X|X \leq x) = \frac{\alpha k}{\lambda F(x)} \sum_{k=1}^{\infty} \sum_{i=0}^{\infty} \Phi_k \frac{(-1)^i}{(i+1)^2} \binom{\alpha k - 1}{i} \{1 - [1 + \lambda(i+1)x] \exp(-\lambda(i+1)x)\}.$$

Hence,

$$B(p) = \frac{\alpha k}{\lambda p \mu} \sum_{k=1}^{\infty} \sum_{i=0}^{\infty} \Phi_k \frac{(-1)^i}{(i+1)^2} \binom{\alpha k - 1}{i} \left\{ 1 - [1 + \lambda(i+1)x_p] \exp(-\lambda(i+1)x_p) \right\}$$

and

$$L(p) = \frac{\alpha k}{\lambda \mu} \sum_{k=1}^{\infty} \sum_{i=0}^{\infty} \Phi_k \frac{(-1)^i}{(i+1)^2} \binom{\alpha k - 1}{i} \left\{ 1 - [1 + \lambda(i+1)x_p] \exp(-\lambda(i+1)x_p) \right\},$$

where  $x_p = -\frac{\log[1-(2^p-1)^{1/\alpha}]}{\lambda}$  is the QF of the WDEx model.

### 3.4. Order statistics

The PDF and CDF of the  $i$ th order statistic for the WDEx distribution are given by

$$\begin{aligned} f_{i:n}(x) &= \frac{f(x)}{B(i, n-i+1)} \sum_{h=0}^{\infty} (-1)^h \binom{n-i}{h} F^{h+i-1}(x) \\ &= \frac{\alpha \lambda n! \log^{-i}(2) (1 - e^{\lambda(-x)})^\alpha \log^{i-1} ((1 - e^{\lambda(-x)})^\alpha + 1) \left(1 - \frac{\log((1 - e^{\lambda(-x)})^\alpha + 1)}{\log(2)}\right)^{n-i}}{\Gamma(i) (e^{\lambda x} - 1) \Gamma(-i+n+1) ((1 - e^{\lambda(-x)})^\alpha + 1)} \end{aligned}$$

and

$$F_{i:n}(x) = \sum_{r=i}^n \binom{n}{r} (F(x))^r (1 - F(x))^{n-r}$$

$$\begin{aligned}
&= \log^{-i}(2) \binom{n}{i} \log^i \left( \left(1 - e^{\lambda(-x)}\right)^{\alpha} + 1 \right) \left( 1 - \frac{\log \left( \left(1 - e^{\lambda(-x)}\right)^{\alpha} + 1 \right)}{\log(2)} \right)^{n-i} \\
&\times {}_2F_1 \left( 1, i-n; i+1; -\frac{\log \left( \left(1 - e^{-x\lambda}\right)^{\alpha} + 1 \right)}{\log \left( \frac{2}{\left(1 - e^{-x\lambda}\right)^{\alpha} + 1} \right)} \right),
\end{aligned}$$

where  ${}_2F_1 \left( 1, i-n; i+1; -\frac{\log \left( \left(1 - e^{-x\lambda}\right)^{\alpha} + 1 \right)}{\log \left( \frac{2}{\left(1 - e^{-x\lambda}\right)^{\alpha} + 1} \right)} \right)$  is a hyper geometric function.

#### 4. Actuarial measures

Five risk measures are explored for the WDEEx distribution in this section. Some actuarial measures such as VaR, TVaR, TV, TVP, and ES are calculated. A brief simulation study for these measures is provided. Finally, an application to the insurance loss data set is analyzed. For more information we refer to refs [17–21].

##### 4.1. VaR measure

The VaR is often referred to as the quantile risk measure or the concept of quantile premium, and it is defined with a certain amount of confidence, say  $q$  (typically 90, 95, or 99%). In comparison, VaR is a quantitative sum of accumulated loss distribution. Risk management are also interested in the “chance of an unfavorable outcome” and can be represented at a specific likelihood level using the VaR. The VaR of a random variable  $X$  is the  $q$ th quantile of its CDF, denoted by  $VaR_q$ , and it is defined by  $VaR_q = Q(q)$  (see Artzner [22]).

The VaR measure of the WDEEx distribution is expressed by

$$VaR_q = - \frac{\log \left[ 1 - (2^q - 1)^{1/\alpha} \right]}{\lambda}.$$

##### 4.2. TVaR measure

The TVaR, which is now being assigned many names, including conditional tail-expectation and conditional VaR, is another significant indicator. The TVaR is used to measure the potential value of the loss provided that there has been an occurrence outside of a provided probability level. The TVaR of a random variable  $X$  is represented by

$$TVaR_q = \frac{1}{(1-q)} \int_{VaR_q}^{\infty} x f(x) dx.$$

The TVaR of the WDEEx distribution is

$$TVaR_q = \frac{\alpha}{(1-q) \lambda \log(2)} \sum_{k=1}^{\infty} \sum_{m=1}^{\infty} \frac{(-1)^{k+1} \left[ 1 - (-1 + 2^q)^{k+\frac{m}{\alpha}} \right]}{m k (m + \alpha k)}. \quad (4.1)$$

#### 4.3. TV measure

Landsman [23] presented the TV risk that is described by the variance of the loss distribution beyond some critical value. The TV of the WDEEx distribution is described as

$$TV_q(X) = E(X^2|X > x_q) - (TVaR_q)^2 = \frac{1}{(1-q)} \int_{VaR_q}^{\infty} x^2 f(x) dx - (TVaR_q)^2, \quad (4.2)$$

where

$$E(X^2|X > x_q) = \frac{2\alpha}{(1-q) \lambda^2 \log(2)} \sum_{k=1}^{\infty} \sum_{m=1}^{\infty} \sum_{n=1}^k \frac{(-1)^{k+1} \left[ 1 - (-1 + 2^q)^{k+\frac{m+1}{\alpha}} \right]}{k n (m+1) (m+\alpha k+1)}, \quad (4.3)$$

by using Eqs (4.1)–(4.3), we get the TV of the WDEEx distribution.

**Table 1.** Simulation results of the five measures of risk for the WDEEx, EEx, and Ex distributions.

Distribution	parameter	Significance level	VaR	TVaR	TV	TVP	ES
WDEEx	$\alpha = 3, \lambda = 0.5$	0.60	3.24353	5.32312	4.07298	7.83746	1.89911
		0.65	3.53435	5.59982	4.08566	8.30964	2.01350
		0.70	3.86421	5.91721	4.09951	8.82180	2.13364
		0.75	4.24850	6.29046	4.11464	9.38883	2.26146
		0.80	4.71278	6.74496	4.13115	10.03667	2.39965
		0.85	5.30475	7.32825	4.14941	10.81284	2.55232
		0.90	6.13120	8.14701	4.16895	11.82411	2.72668
		0.95	7.53258	9.54158	4.19058	13.41092	2.93856
		0.60	0.49017	2.04260	3.25950	3.99830	0.11832
EEx	$\alpha = 3, \lambda = 0.5$	0.65	0.64444	2.25379	3.35743	4.43611	0.15262
		0.70	0.84081	2.50632	3.45651	4.92587	0.19448
		0.75	1.09436	2.81518	3.55661	5.48264	0.24565
		0.80	1.43025	3.20549	3.65760	6.13157	0.30867
		0.85	1.89560	3.72427	3.75935	6.91972	0.38752
		0.90	2.59602	4.47711	3.86176	7.95269	0.48924
		0.95	3.86842	5.80090	3.96470	9.56736	0.62946
		0.60	1.83505	3.83774	4.05243	6.26920	0.77933
		0.65	2.10247	4.10516	4.05243	6.73924	0.87059
Ex	$\lambda = 0.5$	0.70	2.41119	4.41388	4.05243	7.25058	0.96933
		0.75	2.77632	4.77901	4.05243	7.81833	1.07725
		0.80	3.22321	5.22590	4.05243	8.46784	1.19689
		0.85	3.79935	5.80204	4.05243	9.24660	1.33222
		0.90	4.61137	6.61406	4.05243	10.26125	1.49032
		0.95	5.99953	8.00222	4.05243	11.85203	1.68693

#### 4.4. TVP measure

The TVP is also another effective criterion in the insurance fields. The TVP of the WDEEx distribution has the form

$$TVP_q(x) = TVaR_q + q \ TV_q, \quad (4.4)$$

where  $0 < q < 1$ . The TVP of the WDEEx distribution follows by replacing Eqs (4.1) and (4.2) in (4.4).

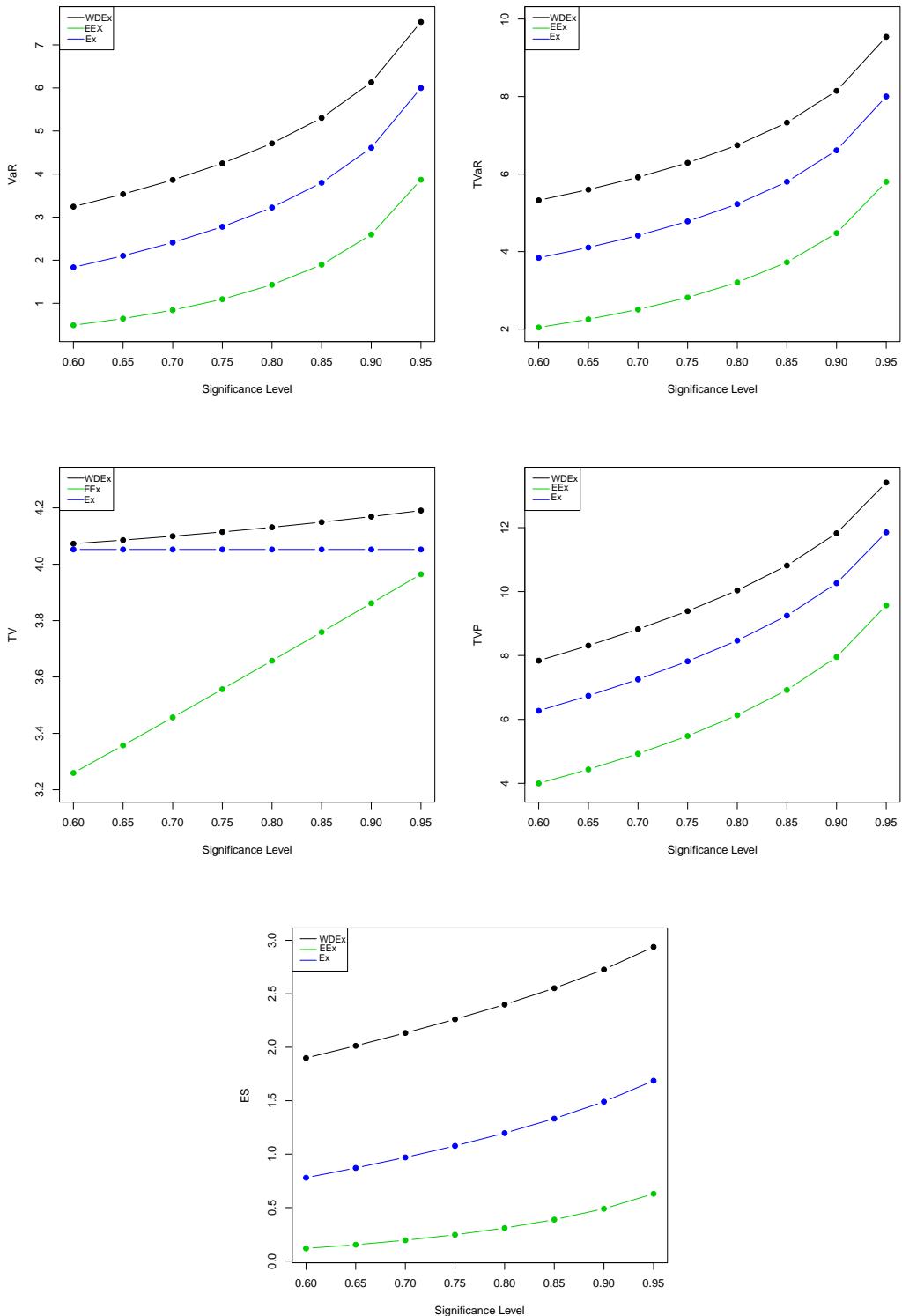
#### 4.5. ES measure

Artzner [22] developed the expected shortfall (ES) as a common indicator in financial risk. In some definitions, if the underlying distribution of  $X$  is continuous then the ES coincides with the Conditional tail expectation and the TVaR. The ES of the WDEEx distribution takes the form

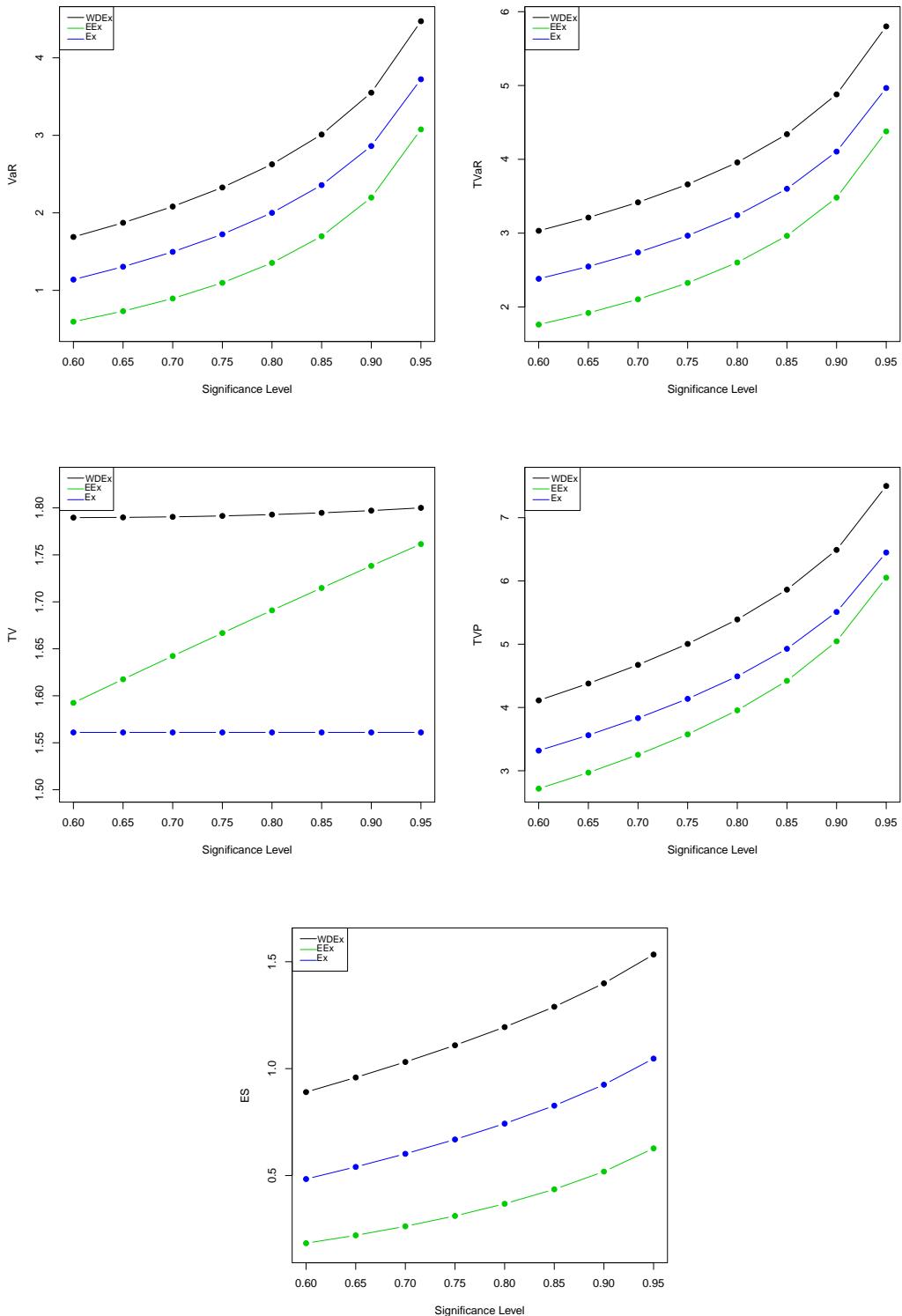
$$ES_q(x) = \frac{1}{q} \int_0^q VaR_t dt = \sum_{m=1}^{\infty} \frac{1}{q m \lambda} \int_0^q (2^t - 1)^{\frac{m}{\alpha}} dt. \quad (4.5)$$

**Table 2.** Simulation results of the five measures of risk for the WDEEx, EEx, and Ex distributions .

Distribution	Parameter	Significance level	VaR	TVaR	TV	TVP	ES
WDEEx	$\alpha = 2, \lambda = 0.75$	0.60	1.68889	3.03197	1.78964	4.11200	0.89027
		0.65	1.87191	3.21103	1.78988	4.37915	0.95860
		0.70	2.08127	3.41720	1.79047	4.67351	1.03113
		0.75	2.32712	3.66052	1.79145	5.00516	1.10910
		0.80	2.62636	3.95782	1.79286	5.39098	1.19423
		0.85	3.01058	4.34064	1.79474	5.86253	1.28924
		0.90	3.55059	4.87971	1.79711	6.49060	1.39887
		0.95	4.47215	5.80076	1.80005	7.50092	1.53359
		0.60	0.59547	1.76160	1.59250	2.71710	0.18354
		0.65	0.73107	1.91880	1.61759	2.97024	0.22029
	EEx	0.70	0.89504	2.10354	1.64237	3.25320	0.26243
		0.75	1.09725	2.32577	1.66681	3.57588	0.31110
		0.80	1.35434	2.60215	1.69096	3.95491	0.36792
		0.85	1.69752	2.96394	1.71477	4.42150	0.43550
		0.90	2.19695	3.48134	1.73829	5.04581	0.51848
		0.95	3.07660	4.37836	1.76151	6.05179	0.62721
Ex	$\lambda = 0.75$	0.60	1.13884	2.38172	1.56099	3.31832	0.48365
		0.65	1.30481	2.54769	1.56099	3.56233	0.54029
		0.70	1.49640	2.73928	1.56100	3.83198	0.60157
		0.75	1.72300	2.96588	1.56099	4.13662	0.66855
		0.80	2.00034	3.24322	1.56099	4.49202	0.74280
		0.85	2.35790	3.60078	1.56099	4.92762	0.82678
		0.90	2.86184	4.10472	1.56099	5.50962	0.92490
		0.95	3.72334	4.96623	1.56099	6.44917	1.04692



**Figure 4.** The above figure displays the graphs of the VaR, TVaR, TV, TVP, and ES using the values in Table 1.



**Figure 5.** The above figure displays the graphs of the VaR, TVaR, TV, TVP, and ES using the values in Table 2.

#### 4.6. Numerical simulations for risk measures

In this sub-section, we discuss some empirical results for the five aforementioned risk measures for the WDEEx, EEx, and Ex distributions for some parametric values. The findings are obtained through the followings steps:

- (1) Random sample of size  $n = 100$  is produced from the WDEEx, EEx, and Ex distributions, and their parameters are calculated via the maximum likelihood (ML) approach.
- (2) 1000 Replications are designed to determine the mean of the VaR, TVaR, TV, TVP as well as ES for the three distributions.

Simulation results of the average VaR, TVaR, TV, TVP and ES for the WDEEx, EEx, and Ex distributions are provided in Tables 1 and 2 and displayed graphically in Figures 4 and 5.

The model with highest value of VaR, TVaR, TV, TVP and ES metrics does also have a heavier tail than other competitor models. The results illustrate that the WDEEx model has improved quantities of the five risk factors than the EEx, and Ex distributions. Thus, the WDEEx model with these parameter values has a heavier-tail than other distributions and could be used efficiently to model heavy-tailed insurance data.

### 5. Classical and Bayesian estimation

This section is devoted to address the estimation of the WDEEx parameters by various estimation methods including eight classical estimation approaches, such as the MLE, Anderson–Darling (ADE) and right-tail Anderson–Darling (RTADE), Cramér–von Mises (CVME), least-squares (LSE), weighted least-squares (WLSE), maximum product of spacings (MPS) and percentile estimation (PCE), and Bayesian approach under three loss functions namely, square error (SE), linear exponential (LINEX), and the general entropy (GE).

Consider the random sample of size  $n$ , say  $x_1, x_2, \dots, x_n$ , from the PDF (2.2), and let  $x_{1:n}, x_{2:n}, \dots, x_{2:n}$  be their associated order statistics.

#### 5.1. Maximum likelihood

The log-likelihood function of the WDEEx distribution has the form

$$L = (\alpha - 1) \sum_{i=1}^n \log((1 - e)^{-\lambda x_i}) - \sum_{i=1}^n \log((1 - e^{-\lambda x_i})^\alpha + 1) - \lambda \sum_{i=1}^n x_i + n \log\left(\frac{\alpha \lambda}{\log(2)}\right). \quad (5.1)$$

Hence, the MLEs of  $\alpha$  and  $\lambda$  are determined by maximizing Eq (5.1) with respect to  $\alpha$  and  $\lambda$ , as the solutions of the following equations

$$\frac{\partial L}{\partial \alpha} = - \sum_{i=1}^n \frac{(1 - e^{-\lambda x_i})^\alpha \log(1 - e^{-\lambda x_i})}{(1 - e^{-\lambda x_i})^\alpha + 1} + \sum_{i=1}^n \log((1 - e)^{-\lambda x_i}) + \frac{n}{\alpha} = 0$$

and

$$\frac{\partial L}{\partial \lambda} = (\alpha - 1) \sum_{i=1}^n \frac{x_i e^{-\lambda x_i}}{1 - e^{-\lambda x_i}} - \sum_{i=1}^n \frac{\alpha x_i e^{-\lambda x_i} (1 - e^{-\lambda x_i})^{\alpha-1}}{(1 - e^{-\lambda x_i})^\alpha + 1} - \sum_{i=1}^n x_i + \frac{n}{\lambda} = 0.$$

Furthermore, solving the above two equations mathematically is complicated, hence the equations can be solved by the numerical methods to obtain the MLE of  $\alpha$  and  $\lambda$ .

### 5.2. Ordinary and weighted least-squares

The LSE of the WDEEx parameters come from minimizing

$$O = \sum_{i=1}^n \left[ F(x_{i:n}) - \frac{i}{n+1} \right]^2 = \sum_{i=1}^n \left[ \frac{\log[(1 - e^{-\lambda x_{i:n}})^\alpha + 1]}{\log(2)} - \frac{i}{n+1} \right]^2.$$

The LSE of the WDEEx parameters are also obtained by solving the following equations

$$\sum_{i=1}^n \left\{ \frac{\log[(1 - e^{-\lambda x_{i:n}})^\alpha + 1]}{\log(2)} - \frac{i}{n+1} \right\} \mathcal{A}_s(x_{i:n}) = 0, \quad s = 1, 2,$$

where

$$\mathcal{A}_1(x_{i:n}) = \frac{\partial}{\partial \alpha} F(x_{i:n}) = \frac{(1 - e^{-\lambda x_i})^\alpha \log(1 - e^{-\lambda x_i})}{\log(2)((1 - e^{-\lambda x_i})^\alpha + 1)}, \quad (5.2)$$

$$\mathcal{A}_2(x_{i:n}) = \frac{\partial}{\partial \lambda} F(x_{i:n}) = \frac{\alpha x_i e^{-\lambda x_i} (1 - e^{-\lambda x_i})^{\alpha-1}}{\log(2)((1 - e^{-\lambda x_i})^\alpha + 1)}. \quad (5.3)$$

The WLSE of the WDEEx parameters can be calculated by minimizing

$$W = \sum_{i=1}^n \frac{(n+1)^2(n+2)}{i(n-i+1)} \left[ F(x_{i:n}) - \frac{i}{n+1} \right]^2 = \sum_{i=1}^n \frac{(n+1)^2(n+2)}{i(n-i+1)} \left[ \frac{\log[(1 - e^{-\lambda x_{i:n}})^\alpha + 1]}{\log(2)} - \frac{i}{n+1} \right]^2.$$

Furthermore, the WLSE of the WDEEx parameters are obtained by solving the following equations

$$\sum_{i=1}^n \frac{(n+1)^2(n+2)}{i(n-i+1)} \left[ \frac{\log[(1 - e^{-\lambda x_{i:n}})^\alpha + 1]}{\log(2)} - \frac{i}{n+1} \right] \mathcal{A}_s(x_{i:n}) = 0,$$

where  $\mathcal{A}_s(x_{i:n}), s = 1, 2, 3$  are given by Eqs (5.2) and (5.3).

### 5.3. Anderson-Darling and right-tail Anderson-Darling

The ADE of the WDEEx parameters are obtained by minimizing

$$A = -n - \frac{1}{n} \sum_{i=1}^n (2i-1)[\log F(x_{i:n}) + \log S(x_{i:n})].$$

The ADE are also calculated by solving the following equations

$$\sum_{i=1}^n (2i-1) \left[ \frac{\mathcal{A}_s(x_{i:n})}{F(x_{i:n})} - \frac{\mathcal{A}_s(x_{n+1-i:n})}{S(x_{n+1-i:n})} \right] = 0,$$

where  $\Delta_s(x_{i:n})$ ,  $s = 1, 2, 3$  were defined in (5.2) and (5.3), respectively.

The RTADE of the WDEx parameters can be calculated by minimizing

$$R = \frac{n}{2} - 2 \sum_{i=1}^n F(x_{i:n}) - \frac{1}{n} \sum_{i=1}^n (2i-1) \log S(x_{i:n}).$$

The RADE can also be follow by solving the non-linear equations

$$-2 \sum_{i=1}^n \Delta_s(x_{n+1-i:n}) + \frac{1}{n} \sum_{i=1}^n \frac{\Delta_s(x_{n+1-i:n})}{S(x_{n+1-i:n})} = 0,$$

where  $\Delta_s(x_{i:n})$ ,  $s = 1, 2, 3$  are defined in (5.2) and (5.3).

#### 5.4. Maximum product of spacings

As an alternative to the ML method, the maximum product spacing (MPS) approach is adopted to estimate the parameters of continuous univariate distributions. The uniform spacings, say  $D_i$ , of a random sample of size  $n$  from the WDEx distribution are defined by

$$D_i = F(x_i) - F(x_{i-1}), \quad i = 1, 2, \dots, n.$$

where  $F(x_0) = 0$ ,  $F(x_{n+1}) = 1$  and  $\sum_{i=1}^{n+1} D_i = 1$ . Then, the MPSE of the WDEx parameters can be obtained by maximizing

$$G = \frac{1}{n+1} \sum_{i=1}^{n+1} \log(D_i),$$

with respect to  $\alpha$  and  $\lambda$ . Further, the MPSE of the WDEx parameters are also obtained by solving

$$\frac{1}{n+1} \sum_{i=1}^{n+1} \frac{1}{D_i} [\Delta_s(x_{i:n}) - \Delta_s(x_{i-1:n})] = 0,$$

where  $\Delta_s(x_{i:n})$ ,  $s = 1, 2, 3$  are presented in (5.2) and (5.3).

#### 5.5. Percentile estimation

Let  $p_i = \frac{i}{n+1}$  be an estimate of  $F(x_{i:n})$ . Hence, the PCE of the WDEx parameters are obtained by minimizing

$$PCE = \sum_{i=1}^n [x_{i:n} - Q(p_i)]^2 = \sum_{i=1}^n \left[ x_{i:n} - \frac{\log [1 - (2^{p_i} - 1)^{1/\alpha}]}{\lambda} \right]^2.$$

The PCE are calculated by solving the following equations

$$\sum_{i=1}^n \left( x_{i:n} - \frac{\log [1 - (2^{p_i} - 1)^{1/\alpha}]}{\lambda} \right) \Phi_s(x_{i:n}) = 0, \quad s = 1, 2,$$

where

$$\begin{aligned} \Phi_1(x_{i:n}) &= \frac{\partial}{\partial \alpha} Q(p_i) = \frac{(2^{p_i} - 1)^{1/\alpha} \log (2^{p_i} - 1)}{\alpha^2 \lambda ((2^{p_i} - 1)^{1/\alpha} - 1)}, \\ \Phi_2(x_{i:n}) &= \frac{\partial}{\partial \lambda} Q(p_i) = \frac{\log (1 - (2^{p_i} - 1)^{1/\alpha})}{\lambda^2}. \end{aligned}$$

### 5.6. Cramér–von mises

The CVME of the WDEx parameters are obtained by minimizing

$$CV = \frac{1}{12n} + \sum_{i=1}^n \left[ F(x_{i:n}) - \frac{2i-1}{2n} \right]^2 = \frac{1}{12n} + \sum_{i=1}^n \left[ \frac{\log[(1-e^{-\lambda x_{i:n}})^\alpha + 1]}{\log(2)} - \frac{2i-1}{2n} \right]^2,$$

They also are obtained by solving the following equations

$$\sum_{i=1}^n \left\{ \frac{\log[(1-e^{-\lambda x_{i:n}})^\alpha + 1]}{\log(2)} - \frac{2i-1}{2n} \right\} \Delta_s(x_{i:n}) = 0,$$

where  $\Delta_s(x_{i:n}), s = 1, 2, 3$  are given in (5.2) and (5.3).

### 5.7. Bayesian estimation

The Bayesian estimation (BE) for the distribution parameters under different loss functions such as the SE, LINEX, and GE loss functions is studied. The parameters are assumed to have gamma priors. Thus,

$$\pi_1(\alpha) \propto \alpha^{\mu_1-1} e^{-\frac{\alpha}{\lambda_1}}, \quad \alpha > 0, \mu_1, \lambda_1 > 0$$

and

$$\pi_2(\lambda) \propto \lambda^{\mu_2-1} e^{-\frac{\lambda}{\lambda_2}}, \quad \lambda > 0, \mu_2, \lambda_2 > 0$$

Suppose that the parameters are independent, therefore the joint PDF of the prior distribution has the form

$$\pi(\alpha, \lambda) \propto \lambda^{\mu_2-1} \alpha^{\mu_1-1} e^{-(\frac{\alpha}{\lambda_1} + \frac{\lambda}{\lambda_2})}. \quad (5.4)$$

The posterior density function of the parameters  $\alpha$  and  $\lambda$  can be obtained from (5.1) and (5.4), as follows

$$\pi^*(\Theta|t) = \pi^*(\alpha, \lambda|t) \propto L(\alpha, \lambda) \pi(\alpha, \lambda). \quad (5.5)$$

The BE of the function of parameters  $U = U(\Theta)$ ,  $\Theta = (\gamma, \alpha, \theta_1, b)$  under the SE loss function (LF) is given by

$$\hat{U}_{SE} = \int_{\Theta} U \pi^*(\Theta|t) d\Theta. \quad (5.6)$$

Under the LINEX LF, the BE of  $U = U(\Theta)$  reduces to

$$\hat{U}_{LINEX} = -\frac{1}{c} \log \left[ \int_{\Theta} e^{-cU} \pi^*(\Theta|t) d\Theta \right], \quad (5.7)$$

where  $c \neq 0$  is the shape parameter of the LINEX LF.

Based on the GE LF, the BE of  $U$  takes the form

$$\hat{U}_{GE} = \left( \int_{\Theta} U^{-q} \pi^*(\Theta|t) d\Theta \right)^{-1/q}, \quad (5.8)$$

where  $q \neq 0$ .

It is obvious that the integrals in Eqs (5.6)–(5.8) are complicated. Consequently, the important sampling method and the MCMC method are applied to obtain an approximation for these integrals. The Metropolis Hasting (MH) algorithm produces a series of draws from the WDE<sub>x</sub> distribution as follows:

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**Algorithm 1** The MCMC for BE

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1. First we will start with initial values for the two parameters  $\alpha^{(0)} = \hat{\alpha}_{MLE}$ ,  $\lambda^{(0)} = \hat{\lambda}_{MLE}$ ,
2. Start the first iteration with counter  $i = 1$ .
3. Now we will generate a random value for  $\alpha^*$  from proposal distribution  $N(\alpha^{(i-1)}, var(\alpha^{(i-1)}))$ , where  $var(\alpha^{(i-1)})$  is the variance of the parameter  $\alpha$
4. We will use the following equation  $A(\alpha^{(i-1)}|\alpha^*) = \min\left[1, \frac{\pi^*(\alpha^*|\lambda^{(i-1)})}{\pi^*(\alpha^{(i-1)}|\lambda^{(i-1)})}\right]$ , the value of this equation denotes the acceptance probability.
5. Now we will generate a random value that follows the uniform distribution  $U \sim U(0, 1)$  that ranges from 0 to 1 .
6. If  $U \leq A(\alpha^{(i-1)}|\alpha^*)$ , put  $\alpha^{(i)} = \alpha^*$ , else put  $\alpha^{(i)} = \alpha^{(i-1)}$ .
7. In order to evaluate the other parameter  $\lambda$  we will repeat the steps from ((2)–(6)) for  $\lambda$ .
8. Now increase the counter by one such that  $i = i + 1$ .
9. To avoid any correlation between the initial values and the estimated values we will repeat the steps from ((3)–(8)),  $N = 10,000$  times to obtain  $(\alpha^{(1)}, \lambda^{(1)}), \dots, (\alpha^{(N)}, \beta^{(N)})$ .

Then, the BEs of  $u(\alpha, \beta)$  using the MCMC under the three LFs are, respectively, given by

$$\tilde{u}_{SE} = \frac{1}{N - M} \sum_{i=M+1}^N u(\alpha^{(i)}, \lambda^{(i)}),$$

$$\tilde{u}_{LINEX} = -\frac{1}{c} \log \left[ \frac{1}{N - M} \sum_{i=M+1}^N \exp\{-c u(\alpha^{(i)}, \lambda^{(i)})\} \right]$$

and

$$\tilde{u}_{GE} = \left[ \frac{1}{N - M} \sum_{i=M+1}^N (u(\alpha^{(i)}, \lambda^{(i)}))^{-q} \right]^{-1/q},$$

where  $M = 2000$  is the burn-in period.

---

### 5.8. TK-approximation for Bayesian estimates

One of the most computationally efficient methods, for estimating the distribution parameter is the Tierney and Kadane [24] (TK) method. They developed an easily calculable integration approximation

approach for evaluating the distribution parameters. The advantage of the TK approach is that it provides an easy and fast approximation method, however some methods are very hard to find the values of integration and takes a lot of time and effort.

Throughout this segment, we use the TK approximation method for calculating the Bayesian estimates (BEs) of the parameters that are focused on the SE, Linex, and GE LFs. TK method [24] is among the techniques to obtain the estimated value of the ratio of the two integrals as given in the following equation

$$\hat{U}(\Theta) = \frac{\int_{\Theta} U(\Theta) e^{[l(\Theta|t)+\rho(\Theta|t)]} d\Theta}{\int_{\Theta} e^{[l(\Theta|t)+\rho(\Theta|t)]} d\Theta}, \quad (5.9)$$

where  $U(\Theta)$  is any function of the parameters  $(\alpha, \lambda)$ ,  $l(\Theta|t)$  is defined in (5.1), and  $\rho(\Theta|t)$  is the logarithm joint prior distribution which is defined by

$$\begin{aligned} \rho(\alpha, \lambda|t) &= (\mu_1 - 1)\ln(\alpha) + (\mu_2 - 1)\ln\lambda \\ &- \left( \frac{\alpha}{\lambda_1} + \frac{\lambda}{\lambda_2} \right). \end{aligned} \quad (5.10)$$

To obtain an explicit expression for  $\hat{U}_{SE}(\Theta)$  using TK approximation, we consider the following functions

$$\delta(\Theta) = \frac{l(\Theta|t) + \rho(\Theta|t)}{n} \quad (5.11)$$

and

$$\delta^*(\Theta) = \delta(\Theta) + \frac{\ln U(\Theta)}{n}. \quad (5.12)$$

Now, assume that the following groups of values  $(\hat{\alpha}_\delta, \hat{\lambda}_\delta)$  and  $(\hat{\alpha}_{\delta^*}, \hat{\lambda}_{\delta^*})$  maximize the functions  $\delta(\alpha, \lambda)$  and  $\delta^*(\alpha, \lambda)$ , respectively.

Hence, we can evaluate the approximation of  $\hat{U}(\Theta)$  by

$$\hat{U}_{TK}(\Theta) = \sqrt{\frac{|\Delta^*|}{|\Delta|}} \exp \left[ n \left\{ \delta(\hat{\alpha}_\delta, \hat{\lambda}_\delta) - \delta(\hat{\alpha}_{\delta^*}, \hat{\lambda}_{\delta^*}) \right\} \right], \quad (5.13)$$

where  $|\Delta|$  and  $|\Delta^*|$  refers to the determinants of negative inverse hessian of  $\delta(\alpha, \lambda)$  and  $\delta^*(\alpha, \lambda)$ , respectively.

## 6. Simulation results

Now, we asses the performance of the above estimation techniques in estimating the WDEx parameters using simulation results. Different sample sizes,  $n = \{15, 30, 70, 100, 200, 350\}$ , and some initial values are assigned. We generated 5000 random samples from the WDEx distribution to calculate the average absolute biases (ABBs), mean square errors (MSEs), and mean relative errors of estimates (MREs) using the R program©.

The following respective mathematical formulae are used to evaluate the ABBs, MSEs, and MREs (for  $\boldsymbol{\theta} = (\alpha, \lambda)'$ )

$$ABs = \frac{1}{N} \sum_{i=1}^N |\widehat{\boldsymbol{\theta}} - \boldsymbol{\theta}|, \quad MSEs = \frac{1}{N} \sum_{i=1}^N (\widehat{\boldsymbol{\theta}} - \boldsymbol{\theta})^2, \quad MREs = \frac{1}{N} \sum_{i=1}^N |\widehat{\boldsymbol{\theta}} - \boldsymbol{\theta}| / \boldsymbol{\theta}.$$

**Table 3.** Simulation results of the WDEx distribution using different estimators for ( $\alpha = 0.75$ ,  $\lambda = 0.75$ ).

n	Est.	Est. Par.	Classical							
			MLE	ADE	CVME	MPS	LSE	PCE	RTADE	WLSE
15	BIAS	$\hat{\alpha}$	0.25192877 <sup>[5][11]</sup>	0.22405917 <sup>[2][8]</sup>	0.30164256 <sup>[7][13]</sup>	0.19839734 <sup>[1][7]</sup>	0.24557181 <sup>[4][10]</sup>	0.32687545 <sup>[8][14]</sup>	0.29638657 <sup>[6][12]</sup>	0.24157906 <sup>[3][9]</sup>
		$\hat{\lambda}$	0.32093518 <sup>[5][11]</sup>	0.30428638 <sup>[2][8]</sup>	0.39524432 <sup>[8][14]</sup>	0.26371696 <sup>[1][7]</sup>	0.31652453 <sup>[4][10]</sup>	0.34677645 <sup>[7][13]</sup>	0.32593945 <sup>[6][12]</sup>	0.31219147 <sup>[3][9]</sup>
	MSE	$\hat{\alpha}$	0.16122813 <sup>[3][9]</sup>	0.11215242 <sup>[2][8]</sup>	0.26608275 <sup>[8][14]</sup>	0.06560425 <sup>[1][7]</sup>	0.1682729 <sup>[4][10]</sup>	0.19013447 <sup>[6][12]</sup>	0.22921014 <sup>[7][13]</sup>	0.18043259 <sup>[5][11]</sup>
		$\hat{\lambda}$	0.21915477 <sup>[4][10]</sup>	0.20639267 <sup>[2][8]</sup>	0.34394628 <sup>[8][14]</sup>	0.11724707 <sup>[1][7]</sup>	0.2237624 <sup>[5][11]</sup>	0.2148123 <sup>[3][9]</sup>	0.23978318 <sup>[7][13]</sup>	0.23922943 <sup>[6][12]</sup>
	MRE	$\hat{\alpha}$	0.33590503 <sup>[5][11]</sup>	0.29874556 <sup>[2][8]</sup>	0.40219008 <sup>[7][13]</sup>	0.26452979 <sup>[1][7]</sup>	0.32742908 <sup>[4][10]</sup>	0.43583393 <sup>[8][14]</sup>	0.3951821 <sup>[6][12]</sup>	0.32210541 <sup>[3][9]</sup>
		$\hat{\lambda}$	0.42791357 <sup>[5][11]</sup>	0.40571517 <sup>[2][8]</sup>	0.52699243 <sup>[8][14]</sup>	0.35162261 <sup>[1][7]</sup>	0.42203271 <sup>[4][10]</sup>	0.4623686 <sup>[7][13]</sup>	0.43458594 <sup>[6][12]</sup>	0.4162553 <sup>[3][9]</sup>
30	$\sum CRanks$		27 <sup>[5]</sup>	12 <sup>[2]</sup>	46 <sup>[8]</sup>	6 <sup>[1]</sup>	25 <sup>[4]</sup>	39 <sup>[7]</sup>	38 <sup>[6]</sup>	23 <sup>[3]</sup>
			63 <sup>[11]</sup>	48 <sup>[8]</sup>	82 <sup>[14]</sup>	42 <sup>[7]</sup>	61 <sup>[10]</sup>	75 <sup>[13]</sup>	74 <sup>[12]</sup>	59 <sup>[9]</sup>
	$\sum C\&BRanks$									
70	BIAS	$\hat{\alpha}$	0.14405974 <sup>[3][9]</sup>	0.13901101 <sup>[2][8]</sup>	0.17725422 <sup>[7][13]</sup>	0.13213658 <sup>[1][7]</sup>	0.16209531 <sup>[5][11]</sup>	0.25820992 <sup>[8][14]</sup>	0.16964029 <sup>[6][12]</sup>	0.14580412 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.18969311 <sup>[2][8]</sup>	0.19141061 <sup>[3][9]</sup>	0.23350009 <sup>[7][13]</sup>	0.18115544 <sup>[1][7]</sup>	0.22786428 <sup>[6][12]</sup>	0.24136118 <sup>[8][14]</sup>	0.20026589 <sup>[4][10]</sup>	0.20766574 <sup>[5][11]</sup>
	MSE	$\hat{\alpha}$	0.03746839 <sup>[3][9]</sup>	0.03635968 <sup>[2][8]</sup>	0.06052504 <sup>[7][13]</sup>	0.02548745 <sup>[1][7]</sup>	0.05095711 <sup>[5][11]</sup>	0.10729773 <sup>[8][14]</sup>	0.05828137 <sup>[6][12]</sup>	0.04260275 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.06746508 <sup>[3][9]</sup>	0.06525692 <sup>[2][8]</sup>	0.09980727 <sup>[8][14]</sup>	0.04907736 <sup>[1][7]</sup>	0.09345408 <sup>[7][13]</sup>	0.09161708 <sup>[6][12]</sup>	0.07713619 <sup>[4][10]</sup>	0.0814058 <sup>[5][11]</sup>
	MRE	$\hat{\alpha}$	0.19207965 <sup>[3][9]</sup>	0.18534802 <sup>[2][8]</sup>	0.23633896 <sup>[7][13]</sup>	0.1761821 <sup>[1][7]</sup>	0.21612709 <sup>[5][11]</sup>	0.34427989 <sup>[8][14]</sup>	0.22618705 <sup>[6][12]</sup>	0.19440549 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.25292415 <sup>[2][8]</sup>	0.25521415 <sup>[3][9]</sup>	0.31133346 <sup>[7][13]</sup>	0.24154058 <sup>[1][7]</sup>	0.30381904 <sup>[6][12]</sup>	0.32181491 <sup>[8][14]</sup>	0.26702119 <sup>[4][10]</sup>	0.27688765 <sup>[5][11]</sup>
$\sum CRanks$			16 <sup>[3]</sup>	14 <sup>[2]</sup>	43 <sup>[7]</sup>	6 <sup>[1]</sup>	34 <sup>[6]</sup>	46 <sup>[8]</sup>	30 <sup>[5]</sup>	27 <sup>[4]</sup>
			52 <sup>[9]</sup>	50 <sup>[8]</sup>	79 <sup>[13]</sup>	42 <sup>[7]</sup>	70 <sup>[12]</sup>	82 <sup>[14]</sup>	66 <sup>[11]</sup>	63 <sup>[10]</sup>
	$\sum C\&BRanks$									
100	BIAS	$\hat{\alpha}$	0.08792199 <sup>[3][9]</sup>	0.08650543 <sup>[2][8]</sup>	0.10310607 <sup>[6][12]</sup>	0.08281633 <sup>[1][7]</sup>	0.10074083 <sup>[5][11]</sup>	0.1753952 <sup>[8][14]</sup>	0.10421227 <sup>[7][13]</sup>	0.08803723 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.11408442 <sup>[1][7]</sup>	0.12445549 <sup>[3][9]</sup>	0.14143312 <sup>[7][13]</sup>	0.1164039 <sup>[2][8]</sup>	0.13839689 <sup>[6][12]</sup>	0.16478766 <sup>[8][14]</sup>	0.12896275 <sup>[5][11]</sup>	0.12549479 <sup>[4][10]</sup>
	MSE	$\hat{\alpha}$	0.01336042 <sup>[3][9]</sup>	0.01302739 <sup>[2][8]</sup>	0.01826087 <sup>[6][12]</sup>	0.01051127 <sup>[1][7]</sup>	0.01740773 <sup>[5][11]</sup>	0.04721909 <sup>[8][14]</sup>	0.01897672 <sup>[7][13]</sup>	0.01339529 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.02286488 <sup>[2][8]</sup>	0.02612201 <sup>[3][9]</sup>	0.03569504 <sup>[7][13]</sup>	0.02027791 <sup>[1][7]</sup>	0.03209949 <sup>[6][12]</sup>	0.04224291 <sup>[8][14]</sup>	0.02824455 <sup>[5][11]</sup>	0.02671826 <sup>[4][10]</sup>
	MRE	$\hat{\alpha}$	0.11722932 <sup>[3][9]</sup>	0.11534057 <sup>[2][8]</sup>	0.13747476 <sup>[6][12]</sup>	0.11042178 <sup>[1][7]</sup>	0.13432111 <sup>[5][11]</sup>	0.23386027 <sup>[8][14]</sup>	0.13894969 <sup>[7][13]</sup>	0.11738297 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.15211257 <sup>[1][7]</sup>	0.16594066 <sup>[3][9]</sup>	0.18857749 <sup>[7][13]</sup>	0.1552052 <sup>[2][8]</sup>	0.18452919 <sup>[6][12]</sup>	0.21971689 <sup>[8][14]</sup>	0.17195033 <sup>[5][11]</sup>	0.16732638 <sup>[4][10]</sup>
$\sum CRanks$			13 <sup>[2]</sup>	15 <sup>[3]</sup>	39 <sup>[7]</sup>	8 <sup>[1]</sup>	33 <sup>[5]</sup>	48 <sup>[8]</sup>	36 <sup>[6]</sup>	24 <sup>[4]</sup>
			49 <sup>[8]</sup>	51 <sup>[9]</sup>	75 <sup>[13]</sup>	44 <sup>[7]</sup>	69 <sup>[11]</sup>	84 <sup>[14]</sup>	72 <sup>[12]</sup>	60 <sup>[10]</sup>
	$\sum C\&BRanks$									
200	BIAS	$\hat{\alpha}$	0.07117082 <sup>[2][8]</sup>	0.07361175 <sup>[3][9]</sup>	0.08687172 <sup>[6][12]</sup>	0.06877471 <sup>[1][7]</sup>	0.07956296 <sup>[5][11]</sup>	0.15707028 <sup>[8][14]</sup>	0.0891824 <sup>[7][13]</sup>	0.07485017 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.10091505 <sup>[4][10]</sup>	0.09696317 <sup>[2][8]</sup>	0.11507711 <sup>[6][12]</sup>	0.0956997 <sup>[1][7]</sup>	0.11614478 <sup>[7][13]</sup>	0.14353138 <sup>[8][14]</sup>	0.10220272 <sup>[5][11]</sup>	0.10079129 <sup>[3][9]</sup>
	MSE	$\hat{\alpha}$	0.00836564 <sup>[2][8]</sup>	0.00880658 <sup>[3][9]</sup>	0.01255586 <sup>[7][13]</sup>	0.0070432 <sup>[1][7]</sup>	0.01016903 <sup>[5][11]</sup>	0.03872841 <sup>[8][14]</sup>	0.01247087 <sup>[6][12]</sup>	0.0095693 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.01716988 <sup>[4][10]</sup>	0.01679548 <sup>[2][8]</sup>	0.02177566 <sup>[6][12]</sup>	0.01380793 <sup>[3][11]</sup>	0.02250852 <sup>[7][13]</sup>	0.03103759 <sup>[8][14]</sup>	0.01734026 <sup>[5][11]</sup>	0.01694889 <sup>[3][9]</sup>
	MRE	$\hat{\alpha}$	0.09489439 <sup>[2][8]</sup>	0.098149 <sup>[3][9]</sup>	0.11582897 <sup>[6][12]</sup>	0.09169961 <sup>[1][7]</sup>	0.10608395 <sup>[5][11]</sup>	0.20942704 <sup>[8][14]</sup>	0.11890987 <sup>[7][13]</sup>	0.09980023 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.13455341 <sup>[4][10]</sup>	0.13284236 <sup>[2][8]</sup>	0.15343615 <sup>[6][12]</sup>	0.1275996 <sup>[1][7]</sup>	0.15485971 <sup>[7][13]</sup>	0.19137517 <sup>[8][14]</sup>	0.1362703 <sup>[5][11]</sup>	0.13438838 <sup>[3][9]</sup>
$\sum CRanks$			18 <sup>[3]</sup>	15 <sup>[2]</sup>	37 <sup>[7]</sup>	6 <sup>[1]</sup>	36 <sup>[6]</sup>	48 <sup>[8]</sup>	35 <sup>[5]</sup>	21 <sup>[4]</sup>
			54 <sup>[9]</sup>	51 <sup>[8]</sup>	73 <sup>[13]</sup>	42 <sup>[7]</sup>	72 <sup>[12]</sup>	84 <sup>[14]</sup>	71 <sup>[11]</sup>	57 <sup>[10]</sup>
	$\sum C\&BRanks$									
300	BIAS	$\hat{\alpha}$	0.03667764 <sup>[1][6]</sup>	0.03990136 <sup>[3][9]</sup>	0.04288476 <sup>[5][11]</sup>	0.03760933 <sup>[2][7]</sup>	0.04356189 <sup>[6][12]</sup>	0.0934606 <sup>[8][14]</sup>	0.04654803 <sup>[7][13]</sup>	0.0400864 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.04795812 <sup>[1][6]</sup>	0.05338798 <sup>[3][9]</sup>	0.05993014 <sup>[6][12]</sup>	0.05257793 <sup>[2][7]</sup>	0.06112134 <sup>[7][13]</sup>	0.07718156 <sup>[8][14]</sup>	0.05407786 <sup>[5][11]</sup>	0.05393484 <sup>[4][10]</sup>
	MSE	$\hat{\alpha}$	0.00215764 <sup>[1][6]</sup>	0.00250001 <sup>[3][8]</sup>	0.00296309 <sup>[6][12]</sup>	0.00221696 <sup>[2][7]</sup>	0.00293099 <sup>[5][11]</sup>	0.01343653 <sup>[8][14]</sup>	0.00346372 <sup>[7][13]</sup>	0.0025318 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.00376997 <sup>[1][6]</sup>	0.00453163 <sup>[3][8]</sup>	0.00573555 <sup>[6][12]</sup>	0.00421602 <sup>[2][7]</sup>	0.00580597 <sup>[7][13]</sup>	0.0091151 <sup>[8][14]</sup>	0.00478321 <sup>[5][10]</sup>	0.00471786 <sup>[4][9]</sup>
	MRE	$\hat{\alpha}$	0.04890353 <sup>[1][6]</sup>	0.05320181 <sup>[3][9]</sup>	0.05717968 <sup>[5][11]</sup>	0.05014577 <sup>[2][7]</sup>	0.05808252 <sup>[6][12]</sup>	0.12461413 <sup>[8][14]</sup>	0.06206403 <sup>[7][13]</sup>	0.053448 <sup>[4][10]</sup>
		$\hat{\lambda}$	0.06394416 <sup>[1][6]</sup>	0.07118398 <sup>[3][9]</sup>	0.07990685 <sup>[6][12]</sup>	0.07010391 <sup>[2][7]</sup>	0.08149512 <sup>[7][13]</sup>	0.10290874 <sup>[8][14]</sup>	0.07210381 <sup>[5][11]</sup>	0.07191312 <sup>[4][10]</sup>
$\sum CRanks$			6 <sup>[1]</sup>	18 <sup>[3]</sup>	34 <sup>[5]</sup>	12 <sup>[2]</sup>	38 <sup>[7]</sup>	48 <sup>[8]</sup>	36 <sup>[6]</sup>	24 <sup>[4]</sup>
			36 <sup>[6]</sup>	52 <sup>[8,5]</sup>	70 <sup>[11]</sup>	42 <sup>[7]</sup>	74 <sup>[13]</sup>	84 <sup>[14]</sup>	71 <sup>[12]</sup>	59 <sup>[10]</sup>
	$\sum C\&BRanks$									

*Continued on next page*

n	Est.	Est. Par.	Bayesian					
			SE	LINEX	GE	SE	LINEX	GE
15	BIAS	$\hat{\alpha}$	0.07080639 <sup>[2][2]</sup>	0.07080648 <sup>[3][3]</sup>	0.07079734 <sup>[1][1]</sup>	0.10887156 <sup>[6][6]</sup>	0.078971 <sup>[5][5]</sup>	0.07215014 <sup>[4][4]</sup>
		$\hat{\lambda}$	0.07334416 <sup>[2][2]</sup>	0.07334414 <sup>[1][1]</sup>	0.07335074 <sup>[3][3]</sup>	0.13371233 <sup>[6][6]</sup>	0.12036804 <sup>[5][5]</sup>	0.11400132 <sup>[4][4]</sup>
	MSE	$\hat{\alpha}$	0.00700589 <sup>[4][4]</sup>	0.0070059 <sup>[5][5]</sup>	0.00700451 <sup>[3][3]</sup>	0.01710565 <sup>[6][6]</sup>	0.00623642 <sup>[2][2]</sup>	0.00520564 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.00734358 <sup>[2][2]</sup>	0.00734357 <sup>[1][1]</sup>	0.00734515 <sup>[3][3]</sup>	0.02521906 <sup>[6][6]</sup>	0.01448847 <sup>[5][5]</sup>	0.0129963 <sup>[4][4]</sup>
	MRE	$\hat{\alpha}$	0.09440852 <sup>[2][2]</sup>	0.09440864 <sup>[3][3]</sup>	0.09439646 <sup>[1][1]</sup>	0.14516208 <sup>[6][6]</sup>	0.10529467 <sup>[5][5]</sup>	0.09620019 <sup>[4][4]</sup>
		$\hat{\lambda}$	0.09779221 <sup>[2][2]</sup>	0.09779219 <sup>[1][1]</sup>	0.09780098 <sup>[3][3]</sup>	0.17828311 <sup>[6][6]</sup>	0.16049072 <sup>[5][5]</sup>	0.15200176 <sup>[4][4]</sup>
	$\sum BRanks$		14 <sup>[2]</sup>	14 <sup>[2]</sup>	14 <sup>[2]</sup>	36 <sup>[6]</sup>	27 <sup>[5]</sup>	21 <sup>[4]</sup>
	$\sum C&BRanks$		14 <sup>[2]</sup>	14 <sup>[2]</sup>	14 <sup>[2]</sup>	36 <sup>[6]</sup>	27 <sup>[5]</sup>	21 <sup>[4]</sup>
30	BIAS	$\hat{\alpha}$	0.06707577 <sup>[4][4]</sup>	0.06707579 <sup>[5][5]</sup>	0.06707479 <sup>[3][3]</sup>	0.08818954 <sup>[6][6]</sup>	0.05398142 <sup>[2][2]</sup>	0.0484456 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.07063728 <sup>[2][2]</sup>	0.07063729 <sup>[3][3]</sup>	0.07063836 <sup>[4][4]</sup>	0.09669995 <sup>[6][6]</sup>	0.07140666 <sup>[5][5]</sup>	0.06590862 <sup>[1][1]</sup>
	MSE	$\hat{\alpha}$	0.00638449 <sup>[4.5][4.5]</sup>	0.00638449 <sup>[4.5][4.5]</sup>	0.00638429 <sup>[3][3]</sup>	0.01167204 <sup>[6][6]</sup>	0.00291399 <sup>[2][2]</sup>	0.00234698 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.00685741 <sup>[4.5][4.5]</sup>	0.00685741 <sup>[4.5][4.5]</sup>	0.00685694 <sup>[3][3]</sup>	0.01388726 <sup>[6][6]</sup>	0.00509891 <sup>[2][2]</sup>	0.00434395 <sup>[1][1]</sup>
	MRE	$\hat{\alpha}$	0.08943436 <sup>[4][4]</sup>	0.08943439 <sup>[5][5]</sup>	0.08943306 <sup>[3][3]</sup>	0.11758605 <sup>[6][6]</sup>	0.07197522 <sup>[2][2]</sup>	0.06459413 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.09418304 <sup>[2][2]</sup>	0.09418306 <sup>[3][3]</sup>	0.09418448 <sup>[4][4]</sup>	0.12893327 <sup>[6][6]</sup>	0.09520888 <sup>[5][5]</sup>	0.08787816 <sup>[1][1]</sup>
	$\sum BRanks$		21 <sup>[4]</sup>	25 <sup>[5]</sup>	20 <sup>[3]</sup>	36 <sup>[6]</sup>	18 <sup>[2]</sup>	6 <sup>[1]</sup>
	$\sum C&BRanks$		21 <sup>[4]</sup>	25 <sup>[5]</sup>	20 <sup>[3]</sup>	36 <sup>[6]</sup>	18 <sup>[2]</sup>	6 <sup>[1]</sup>
70	BIAS	$\hat{\alpha}$	0.0628967 <sup>[4][4]</sup>	0.06289672 <sup>[5][5]</sup>	0.06289578 <sup>[3][3]</sup>	0.06988905 <sup>[6][6]</sup>	0.03512338 <sup>[2][2]</sup>	0.03097589 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.06382592 <sup>[3][3]</sup>	0.06382593 <sup>[4][4]</sup>	0.06382623 <sup>[5][5]</sup>	0.08500664 <sup>[6][6]</sup>	0.04706336 <sup>[2][2]</sup>	0.04163393 <sup>[1][1]</sup>
	MSE	$\hat{\alpha}$	0.00529453 <sup>[4.5][4.5]</sup>	0.00529453 <sup>[4.5][4.5]</sup>	0.00529439 <sup>[3][3]</sup>	0.00766215 <sup>[6][6]</sup>	0.00123365 <sup>[2][2]</sup>	0.00095951 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.00577517 <sup>[3.5][3.5]</sup>	0.00577517 <sup>[3.5][3.5]</sup>	0.00577528 <sup>[5][5]</sup>	0.01076173 <sup>[6][6]</sup>	0.00221496 <sup>[2][2]</sup>	0.00173338 <sup>[1][1]</sup>
	MRE	$\hat{\alpha}$	0.08386227 <sup>[4][4]</sup>	0.08386229 <sup>[5][5]</sup>	0.08386104 <sup>[3][3]</sup>	0.0931854 <sup>[6][6]</sup>	0.04683117 <sup>[2][2]</sup>	0.04130118 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.08510123 <sup>[3.5][3.5]</sup>	0.08510123 <sup>[3.5][3.5]</sup>	0.08510164 <sup>[5][5]</sup>	0.11334219 <sup>[6][6]</sup>	0.06275115 <sup>[2][2]</sup>	0.05551119 <sup>[1][1]</sup>
	$\sum BRanks$		22.5 <sup>[3]</sup>	25.5 <sup>[5]</sup>	24 <sup>[4]</sup>	36 <sup>[6]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>
	$\sum C&BRanks$		22.5 <sup>[3]</sup>	25.5 <sup>[5]</sup>	24 <sup>[4]</sup>	36 <sup>[6]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>
100	BIAS	$\hat{\alpha}$	0.05470376 <sup>[4][4]</sup>	0.05470375 <sup>[3][3]</sup>	0.05470411 <sup>[5][5]</sup>	0.06296533 <sup>[6][6]</sup>	0.02807831 <sup>[2][2]</sup>	0.02458989 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.06312911 <sup>[3.5][3.5]</sup>	0.06312911 <sup>[3.5][3.5]</sup>	0.0631295 <sup>[5][5]</sup>	0.07802856 <sup>[6][6]</sup>	0.04616373 <sup>[2][2]</sup>	0.04172589 <sup>[1][1]</sup>
	MSE	$\hat{\alpha}$	0.00439471 <sup>[3.5][3.5]</sup>	0.00439471 <sup>[3.5][3.5]</sup>	0.00439478 <sup>[5][5]</sup>	0.0608088 <sup>[6][6]</sup>	0.00078839 <sup>[2][2]</sup>	0.00060466 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.00567389 <sup>[4.5][4.5]</sup>	0.00567389 <sup>[4.5][4.5]</sup>	0.00567383 <sup>[3][3]</sup>	0.00900842 <sup>[6][6]</sup>	0.00213109 <sup>[2][2]</sup>	0.00174105 <sup>[1][1]</sup>
	MRE	$\hat{\alpha}$	0.07293834 <sup>[3.5][3.5]</sup>	0.07293834 <sup>[3.5][3.5]</sup>	0.07293819 <sup>[5][5]</sup>	0.08395377 <sup>[6][6]</sup>	0.03743775 <sup>[2][2]</sup>	0.03278652 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.08417215 <sup>[3.5][3.5]</sup>	0.08417215 <sup>[3.5][3.5]</sup>	0.08417267 <sup>[5][5]</sup>	0.10403808 <sup>[6][6]</sup>	0.06155164 <sup>[2][2]</sup>	0.05563452 <sup>[1][1]</sup>
	$\sum BRanks$		22.5 <sup>[4]</sup>	21.5 <sup>[3]</sup>	28 <sup>[5]</sup>	36 <sup>[6]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>
	$\sum C&BRanks$		22.5 <sup>[4]</sup>	21.5 <sup>[3]</sup>	28 <sup>[5]</sup>	36 <sup>[6]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>
200	BIAS	$\hat{\alpha}$	0.04001852 <sup>[4.5][4.5]</sup>	0.04001852 <sup>[4.5][4.5]</sup>	0.04001848 <sup>[3][3]</sup>	0.04705698 <sup>[6][6]</sup>	0.01733801 <sup>[2][2]</sup>	0.01516892 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.05856409 <sup>[4][4]</sup>	0.0585641 <sup>[5][5]</sup>	0.05856347 <sup>[3][3]</sup>	0.06345567 <sup>[6][6]</sup>	0.02796569 <sup>[2][2]</sup>	0.02441503 <sup>[1][1]</sup>
	MSE	$\hat{\alpha}$	0.00253077 <sup>[4.5][4.5]</sup>	0.00253077 <sup>[4.5][4.5]</sup>	0.00253076 <sup>[3][3]</sup>	0.00362108 <sup>[6][6]</sup>	0.00030061 <sup>[2][2]</sup>	0.0002301 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.00476338 <sup>[4.5][4.5]</sup>	0.00476338 <sup>[4.5][4.5]</sup>	0.004763 <sup>[3][3]</sup>	0.00626275 <sup>[6][6]</sup>	0.00078208 <sup>[2][2]</sup>	0.00059609 <sup>[1][1]</sup>
	MRE	$\hat{\alpha}$	0.05335802 <sup>[4.5][4.5]</sup>	0.05335802 <sup>[4.5][4.5]</sup>	0.05335798 <sup>[3][3]</sup>	0.06274264 <sup>[6][6]</sup>	0.02311735 <sup>[2][2]</sup>	0.02022522 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.07808545 <sup>[4][4]</sup>	0.07808546 <sup>[5][5]</sup>	0.07808462 <sup>[3][3]</sup>	0.08460756 <sup>[6][6]</sup>	0.03728759 <sup>[2][2]</sup>	0.03255338 <sup>[1][1]</sup>
	$\sum BRanks$		26 <sup>[4]</sup>	28 <sup>[5]</sup>	18 <sup>[3]</sup>	36 <sup>[6]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>
	$\sum C&BRanks$		26 <sup>[4]</sup>	28 <sup>[5]</sup>	18 <sup>[3]</sup>	36 <sup>[6]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>
300	BIAS	$\hat{\alpha}$	0.03245976 <sup>[4.5][4.5]</sup>	0.03245976 <sup>[4.5][4.5]</sup>	0.03245972 <sup>[3][3]</sup>	0.03938224 <sup>[6][8]</sup>	0.01407489 <sup>[2][2]</sup>	0.01252037 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.04732128 <sup>[3.5][3.5]</sup>	0.04732128 <sup>[3.5][3.5]</sup>	0.0473218 <sup>[5][5]</sup>	0.05290774 <sup>[6][8]</sup>	0.02365826 <sup>[2][2]</sup>	0.02092282 <sup>[1][1]</sup>
	MSE	$\hat{\alpha}$	0.00166778 <sup>[4][4]</sup>	0.00166778 <sup>[4][4]</sup>	0.00166778 <sup>[4][4]</sup>	0.00252766 <sup>[6][9]</sup>	0.0001981 <sup>[2][2]</sup>	0.00015676 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.00341064 <sup>[3.5][3.5]</sup>	0.00341064 <sup>[3.5][3.5]</sup>	0.00341069 <sup>[5][5]</sup>	0.00482747 <sup>[6][11]</sup>	0.00055971 <sup>[2][2]</sup>	0.00043776 <sup>[1][1]</sup>
	MRE	$\hat{\alpha}$	0.04327968 <sup>[4.5][4.5]</sup>	0.04327968 <sup>[4.5][4.5]</sup>	0.04327963 <sup>[3][3]</sup>	0.05250965 <sup>[6][8]</sup>	0.01876651 <sup>[2][2]</sup>	0.01669383 <sup>[1][1]</sup>
		$\hat{\lambda}$	0.06309505 <sup>[4][4]</sup>	0.06309504 <sup>[3][3]</sup>	0.06309574 <sup>[5][5]</sup>	0.07054365 <sup>[6][8]</sup>	0.03154434 <sup>[2][2]</sup>	0.02789709 <sup>[1][1]</sup>
	$\sum BRanks$		24 <sup>[4]</sup>	23 <sup>[3]</sup>	25 <sup>[5]</sup>	36 <sup>[6]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>
	$\sum C&BRanks$		24 <sup>[4]</sup>	23 <sup>[3]</sup>	25 <sup>[5]</sup>	52 <sup>[8.5]</sup>	12 <sup>[2]</sup>	6 <sup>[1]</sup>

Table 3 reports the simulation results including the ABBs, MSEs, and MREs of the WDE<sub>x</sub> parameters using the nine estimation approaches. One can note that all methods provide small ABBs, MSEs and MREs in all parametric values. The nine estimators are consistent, i.e., the MSEs decrease as the sample size increases, for all studied cases.

**Table 4.** The CI length for different parametric values of all estimation methods.

n	Est. Par.	MLE			ADE			CVME			MPS			LSE			PCE			RTADE			WLSE		
		L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. BP	L. BT	L. CR			
15	$\hat{\alpha}$	1.76864 <sup>[4]</sup>	1.01231 <sup>[5]</sup>	1.23169 <sup>[7]</sup>	0.9917 <sup>[4]</sup>	2.44613 <sup>[6]</sup>	0.75104 <sup>[1]</sup>	1.42478 <sup>[11]</sup>	0.85626 <sup>[3]</sup>	1.37582 <sup>[10]</sup>	1.2964 <sup>[9]</sup>	1.73329 <sup>[13]</sup>	4.52643 <sup>[17]</sup>	2.21266 <sup>[15]</sup>	1.4343 <sup>[12]</sup>	1.26902 <sup>[8]</sup>	1.09219 <sup>[6]</sup>	0.79185 <sup>[2]</sup>							
	$\hat{\lambda}$	1.8179 <sup>[3]</sup>	1.21276 <sup>[3]</sup>	1.61918 <sup>[9]</sup>	1.35754 <sup>[6]</sup>	2.45476 <sup>[6]</sup>	1.76051 <sup>[12]</sup>	1.08184 <sup>[2]</sup>	1.21353 <sup>[4]</sup>	1.82447 <sup>[14]</sup>	1.7064 <sup>[10]</sup>	1.51256 <sup>[7]</sup>	2.71437 <sup>[17]</sup>	1.85662 <sup>[15]</sup>	1.32039 <sup>[5]</sup>	1.73185 <sup>[11]</sup>	1.56473 <sup>[8]</sup>	1.02613 <sup>[1]</sup>							
30	$\hat{\alpha}$	0.76771 <sup>[10]</sup>	0.61941 <sup>[3]</sup>	0.69962 <sup>[6]</sup>	0.63344 <sup>[5]</sup>	0.93483 <sup>[4]</sup>	0.73927 <sup>[8]</sup>	0.54674 <sup>[1]</sup>	0.62416 <sup>[4]</sup>	0.8346 <sup>[13]</sup>	0.82046 <sup>[12]</sup>	1.17514 <sup>[16]</sup>	2.60316 <sup>[17]</sup>	0.96934 <sup>[15]</sup>	0.81431 <sup>[11]</sup>	0.75708 <sup>[9]</sup>	0.73181 <sup>[7]</sup>	0.56731 <sup>[2]</sup>							
	$\hat{\lambda}$	1.01832 <sup>[9]</sup>	0.85197 <sup>[4]</sup>	0.95639 <sup>[7]</sup>	0.87947 <sup>[6]</sup>	1.2636 <sup>[16]</sup>	1.08914 <sup>[13]</sup>	0.74985 <sup>[2]</sup>	0.85211 <sup>[5]</sup>	1.15014 <sup>[15]</sup>	1.14232 <sup>[14]</sup>	1.0211 <sup>[10]</sup>	1.58731 <sup>[17]</sup>	1.04276 <sup>[12]</sup>	0.81825 <sup>[5]</sup>	1.03151 <sup>[11]</sup>	1.00314 <sup>[8]</sup>	0.73465 <sup>[1]</sup>							
70	$\hat{\alpha}$	0.44252 <sup>[9]</sup>	0.40638 <sup>[3]</sup>	0.43855 <sup>[8]</sup>	0.4246 <sup>[5]</sup>	0.52287 <sup>[15]</sup>	0.47767 <sup>[11]</sup>	0.3631 <sup>[1]</sup>	0.4313 <sup>[4]</sup>	0.47865 <sup>[12]</sup>	0.47271 <sup>[10]</sup>	0.82108 <sup>[6]</sup>	1.36637 <sup>[17]</sup>	0.51984 <sup>[14]</sup>	0.47789 <sup>[13]</sup>	0.43536 <sup>[6]</sup>	0.43582 <sup>[7]</sup>	0.37696 <sup>[2]</sup>							
	$\hat{\lambda}$	0.59365 <sup>[6]</sup>	0.54695 <sup>[4]</sup>	0.61226 <sup>[10]</sup>	0.59699 <sup>[7]</sup>	0.75869 <sup>[16]</sup>	0.69736 <sup>[14]</sup>	0.4837 <sup>[1]</sup>	0.54884 <sup>[5]</sup>	0.65062 <sup>[12]</sup>	0.6587 <sup>[13]</sup>	0.70313 <sup>[15]</sup>	0.80559 <sup>[17]</sup>	0.60571 <sup>[8]</sup>	0.50151 <sup>[3]</sup>	0.60658 <sup>[9]</sup>	0.61611 <sup>[11]</sup>	0.49915 <sup>[2]</sup>							
100	$\hat{\alpha}$	0.39444 <sup>[5]</sup>	0.31992 <sup>[3]</sup>	0.34198 <sup>[6]</sup>	0.33375 <sup>[4]</sup>	0.442215 <sup>[14]</sup>	0.38962 <sup>[10]</sup>	0.30229 <sup>[1]</sup>	0.34622 <sup>[7]</sup>	0.39767 <sup>[11]</sup>	0.39816 <sup>[12]</sup>	0.68273 <sup>[16]</sup>	1.05362 <sup>[17]</sup>	0.43551 <sup>[5]</sup>	0.41384 <sup>[13]</sup>	0.3649 <sup>[9]</sup>	0.35835 <sup>[8]</sup>	0.31696 <sup>[2]</sup>							
	$\hat{\lambda}$	0.47339 <sup>[6]</sup>	0.43589 <sup>[4]</sup>	0.47704 <sup>[7]</sup>	0.47049 <sup>[5]</sup>	0.56073 <sup>[15]</sup>	0.52885 <sup>[12]</sup>	0.40649 <sup>[11]</sup>	0.4866 <sup>[8]</sup>	0.54784 <sup>[13]</sup>	0.55029 <sup>[14]</sup>	0.59234 <sup>[16]</sup>	0.62163 <sup>[17]</sup>	0.52259 <sup>[11]</sup>	0.43411 <sup>[3]</sup>	0.51049 <sup>[10]</sup>	0.5069 <sup>[9]</sup>	0.42203 <sup>[2]</sup>							
200	$\hat{\alpha}$	0.2389 <sup>[4]</sup>	0.23142 <sup>[3]</sup>	0.2516 <sup>[9]</sup>	0.2482 <sup>[8]</sup>	0.27946 <sup>[13]</sup>	0.27005 <sup>[10]</sup>	0.22294 <sup>[11]</sup>	0.24721 <sup>[7]</sup>	0.27286 <sup>[12]</sup>	0.27282 <sup>[11]</sup>	0.52575 <sup>[16]</sup>	0.70941 <sup>[17]</sup>	0.29224 <sup>[15]</sup>	0.2862 <sup>[14]</sup>	0.24663 <sup>[6]</sup>	0.2461 <sup>[5]</sup>	0.2278 <sup>[2]</sup>							
	$\hat{\lambda}$	0.32208 <sup>[5]</sup>	0.31305 <sup>[4]</sup>	0.35705 <sup>[11]</sup>	0.3489 <sup>[9]</sup>	0.3895 <sup>[5]</sup>	0.37969 <sup>[12]</sup>	0.30621 <sup>[2]</sup>	0.34766 <sup>[8]</sup>	0.38153 <sup>[14]</sup>	0.3812 <sup>[13]</sup>	0.43553 <sup>[17]</sup>	0.39775 <sup>[16]</sup>	0.35301 <sup>[10]</sup>	0.30882 <sup>[3]</sup>	0.33943 <sup>[6]</sup>	0.34042 <sup>[7]</sup>	0.3042 <sup>[11]</sup>							
350	$\hat{\alpha}$	0.17441 <sup>[4]</sup>	0.17146 <sup>[2]</sup>	0.1858 <sup>[6]</sup>	0.18336 <sup>[5]</sup>	0.20941 <sup>[13]</sup>	0.20323 <sup>[9]</sup>	0.17002 <sup>[11]</sup>	0.21806 <sup>[15]</sup>	0.20592 <sup>[10]</sup>	0.20694 <sup>[11]</sup>	0.40697 <sup>[16]</sup>	0.49775 <sup>[17]</sup>	0.21169 <sup>[14]</sup>	0.20828 <sup>[12]</sup>	0.19022 <sup>[7]</sup>	0.19045 <sup>[8]</sup>	0.17249 <sup>[3]</sup>							
	$\hat{\lambda}$	0.24024 <sup>[5]</sup>	0.23525 <sup>[4]</sup>	0.25748 <sup>[8]</sup>	0.2569 <sup>[7]</sup>	0.29707 <sup>[15]</sup>	0.28529 <sup>[12]</sup>	0.23099 <sup>[2]</sup>	0.30366 <sup>[16]</sup>	0.28776 <sup>[13]</sup>	0.29495 <sup>[14]</sup>	0.34137 <sup>[17]</sup>	0.27981 <sup>[11]</sup>	0.25913 <sup>[10]</sup>	0.22713 <sup>[11]</sup>	0.25778 <sup>[9]</sup>	0.25635 <sup>[6]</sup>	0.23339 <sup>[3]</sup>							
$\sum Rank_s$		90 <sup>[6,5]</sup>	42 <sup>[3]</sup>	94 <sup>[9]</sup>	71 <sup>[4]</sup>	178 <sup>[16]</sup>	134 <sup>[11]</sup>	16 <sup>[1]</sup>	86 <sup>[5]</sup>	149 <sup>[3]</sup>	143 <sup>[12]</sup>	175 <sup>[15]</sup>	197 <sup>[17]</sup>	154 <sup>[4]</sup>	93 <sup>[8]</sup>	101 <sup>[10]</sup>	90 <sup>[6,5]</sup>	23 <sup>[2]</sup>							

**Table 5.** Overall probabilities for all estimation methods .

n	Est. Par.	MLE		ADE		CVME		MPS		LSE		PCE		RTADE		WLSE		Bayesian
		CV. BP	CV. BT	CV. BP	CV. BT	CV. BP	CV. BT	CV. BP	CV. BT	CV. BP	CV. BT	CV. BP	CV. BT	CV. BP	CV. BT	CV. BP	CV. BT	
15	$\hat{\alpha}$	0.8 <sup>{10}</sup>	0.85 <sup>{3}</sup>	0.93 <sup>{10.5}</sup>	0.95 <sup>{10.6}</sup>	0.84 <sup>{2}</sup>	0.9 <sup>{6}</sup>	0.91 <sup>{8}</sup>	0.94 <sup>{13.5}</sup>	0.93 <sup>{10.5}</sup>	0.92 <sup>{9}</sup>	0.86 <sup>{4}</sup>	0.9 <sup>{6}</sup>	0.94 <sup>{13.5}</sup>	0.94 <sup>{13.5}</sup>	0.94 <sup>{13.5}</sup>	0.94 <sup>{13.5}</sup>	1 <sup>{17}</sup>
	$\lambda$	0.82 <sup>{15}</sup>	0.89 <sup>{7}</sup>	0.94 <sup>{14.5}</sup>	0.91 <sup>{9}</sup>	0.85 <sup>{4}</sup>	0.88 <sup>{6}</sup>	0.93 <sup>{11.5}</sup>	0.97 <sup>{16}</sup>	0.94 <sup>{14.5}</sup>	0.9 <sup>{8}</sup>	0.93 <sup>{11.5}</sup>	0.84 <sup>{3}</sup>	0.93 <sup>{11.5}</sup>	0.82 <sup>{15}</sup>	0.93 <sup>{11.5}</sup>	0.87 <sup>{5}</sup>	1 <sup>{17}</sup>
30	$\hat{\alpha}$	0.88 <sup>{4}</sup>	0.92 <sup>{11}</sup>	0.94 <sup>{13.5}</sup>	0.88 <sup>{4}</sup>	0.92 <sup>{11}</sup>	0.91 <sup>{8.5}</sup>	0.85 <sup>{11}</sup>	0.88 <sup>{4}</sup>	0.89 <sup>{6}</sup>	0.91 <sup>{8.5}</sup>	0.86 <sup>{2}</sup>	0.9 <sup>{7}</sup>	0.97 <sup>{16}</sup>	0.94 <sup>{13.5}</sup>	0.95 <sup>{15}</sup>	0.92 <sup>{11}</sup>	1 <sup>{17}</sup>
	$\lambda$	0.89 <sup>{4}</sup>	0.92 <sup>{8.5}</sup>	0.98 <sup>{6}</sup>	0.91 <sup>{6}</sup>	0.93 <sup>{11}</sup>	0.92 <sup>{8.5}</sup>	0.91 <sup>{6}</sup>	0.94 <sup>{13}</sup>	0.95 <sup>{14.5}</sup>	0.93 <sup>{11}</sup>	0.87 <sup>{3}</sup>	0.86 <sup>{2}</sup>	0.91 <sup>{6}</sup>	0.81 <sup>{11}</sup>	0.95 <sup>{14.5}</sup>	0.93 <sup>{11}</sup>	1 <sup>{17}</sup>
70	$\hat{\alpha}$	0.94 <sup>{9}</sup>	0.97 <sup>{16}</sup>	0.91 <sup>{4.5}</sup>	0.95 <sup>{12.5}</sup>	0.89 <sup>{2.5}</sup>	0.89 <sup>{2.5}</sup>	0.84 <sup>{11}</sup>	0.94 <sup>{9}</sup>	0.93 <sup>{6.5}</sup>	0.95 <sup>{12.5}</sup>	0.93 <sup>{6.5}</sup>	0.95 <sup>{12.5}</sup>	0.95 <sup>{12.5}</sup>	0.91 <sup>{4.5}</sup>	0.94 <sup>{9}</sup>	0.96 <sup>{15}</sup>	0.978 <sup>{17}</sup>
	$\lambda$	0.87 <sup>{2.5}</sup>	0.93 <sup>{12}</sup>	0.93 <sup>{12}</sup>	0.92 <sup>{10}</sup>	0.91 <sup>{8.5}</sup>	0.87 <sup>{2.5}</sup>	0.88 <sup>{4.5}</sup>	0.88 <sup>{4.5}</sup>	0.89 <sup>{13}</sup>	0.91 <sup>{8.5}</sup>	0.94 <sup>{14}</sup>	0.9 <sup>{6.5}</sup>	0.93 <sup>{12}</sup>	0.83 <sup>{11}</sup>	0.95 <sup>{15.5}</sup>	0.95 <sup>{15.5}</sup>	0.993 <sup>{17}</sup>
100	$\hat{\alpha}$	0.94 <sup>{10.5}</sup>	0.96 <sup>{13}</sup>	0.94 <sup>{10.5}</sup>	0.92 <sup>{5.5}</sup>	0.9 <sup>{3}</sup>	0.96 <sup>{13}</sup>	0.82 <sup>{11}</sup>	0.93 <sup>{8}</sup>	0.96 <sup>{13}</sup>	0.97 <sup>{16.5}</sup>	0.89 <sup>{2}</sup>	0.92 <sup>{5.5}</sup>	0.97 <sup>{16.5}</sup>	0.93 <sup>{8}</sup>	0.93 <sup>{8}</sup>	0.91 <sup>{4}</sup>	0.967 <sup>{15}</sup>
	$\lambda$	0.93 <sup>{12}</sup>	0.91 <sup>{8.5}</sup>	0.96 <sup>{13}</sup>	0.96 <sup>{10}</sup>	0.96 <sup>{16}</sup>	0.93 <sup>{12}</sup>	0.86 <sup>{2}</sup>	0.9 <sup>{6}</sup>	0.95 <sup>{15}</sup>	0.91 <sup>{8.5}</sup>	0.97 <sup>{16.5}</sup>	0.89 <sup>{2}</sup>	0.92 <sup>{5.5}</sup>	0.93 <sup>{12}</sup>	0.88 <sup>{3.5}</sup>	0.93 <sup>{12}</sup>	0.964 <sup>{17}</sup>
200	$\hat{\alpha}$	0.93 <sup>{9.5}</sup>	0.95 <sup>{15}</sup>	0.95 <sup>{15}</sup>	0.91 <sup>{6.5}</sup>	0.94 <sup>{12.5}</sup>	0.93 <sup>{9.5}</sup>	0.93 <sup>{9.5}</sup>	0.94 <sup>{4.5}</sup>	0.94 <sup>{4.5}</sup>	0.91 <sup>{6.5}</sup>	0.85 <sup>{11}</sup>	0.95 <sup>{15}</sup>	0.94 <sup>{12.5}</sup>	0.93 <sup>{9.5}</sup>	0.88 <sup>{3}</sup>	0.96 <sup>{2}</sup>	0.984 <sup>{17}</sup>
	$\lambda$	0.9 <sup>{6}</sup>	0.88 <sup>{5}</sup>	0.95 <sup>{14}</sup>	0.93 <sup>{11.5}</sup>	0.96 <sup>{15}</sup>	0.97 <sup>{16}</sup>	0.87 <sup>{3.5}</sup>	0.91 <sup>{7.5}</sup>	0.93 <sup>{11.5}</sup>	0.91 <sup>{7.5}</sup>	0.86 <sup>{2}</sup>	0.81 <sup>{11}</sup>	0.92 <sup>{9.5}</sup>	0.87 <sup>{3.5}</sup>	0.98 <sup>{17}</sup>	0.972 <sup>{17}</sup>	
350	$\hat{\alpha}$	0.92 <sup>{9}</sup>	0.91 <sup>{7}</sup>	0.89 <sup>{4.5}</sup>	0.89 <sup>{4.5}</sup>	0.94 <sup>{13}</sup>	0.91 <sup>{7}</sup>	0.93 <sup>{11}</sup>	0.93 <sup>{11}</sup>	0.98 <sup>{17}</sup>	0.97 <sup>{16}</sup>	0.87 <sup>{11}</sup>	0.96 <sup>{14}</sup>	0.92 <sup>{9.5}</sup>	0.98 <sup>{17}</sup>	0.92 <sup>{9.5}</sup>	0.949 <sup>{13}</sup>	
	$\lambda$	0.95 <sup>{15}</sup>	0.97 <sup>{17}</sup>	0.92 <sup>{5}</sup>	0.94 <sup>{11}</sup>	0.94 <sup>{11}</sup>	0.91 <sup>{3}</sup>	0.92 <sup>{5}</sup>	0.94 <sup>{11}</sup>	0.95 <sup>{15}</sup>	0.92 <sup>{5}</sup>	0.83 <sup>{11}</sup>	0.93 <sup>{7}</sup>	0.88 <sup>{2}</sup>	0.94 <sup>{11}</sup>	0.95 <sup>{15}</sup>	0.936 <sup>{8}</sup>	
$\sum Ranks$		84 <sup>{5}</sup>	123 <sup>{11}</sup>	126 <sup>{13}</sup>	112.5 <sup>{9}</sup>	105.5 <sup>{8}</sup>	102.5 <sup>{7}</sup>	62 <sup>{3}</sup>	102 <sup>{6}</sup>	135.5 <sup>{5}</sup>	127 <sup>{14}</sup>	60.5 <sup>{2}</sup>	72.5 <sup>{4}</sup>	124 <sup>{12}</sup>	56.5 <sup>{11}</sup>	141 <sup>{16}</sup>	114.5 <sup>{10}</sup>	187 <sup>{17}</sup>

## 7. Modeling an insurance data

Here, we analyze a data set from the insurance sector to illustrate the utility of the WDEx distribution. This data set reflects the monthly unemployment insurance metrics from July 2008 to April 2013, and it includes 58 observations as recorded by the department of labor, state of Maryland, USA. The data frame consists of 21 variables and the variable number 9 is fitted using the WDEx model. Data is listed below and it can be found at: <https://catalog.data.gov/dataset/unemployment-insurance-data-from-July-2008-to-April-2013>.

48.2	48.2	55.6	49.4	54.7	84.4	85.6	92.1	115.7	92.7
87.4	104.1	86.6	96.7	74.1	70.4	84	80.3	86.5	86.9
104.7	74.3	66.7	79.1	65.7	78.6	59.9	59	72.6	66.6
89.9	75.2	72.1	63.2	70.3	58.9	59.7	72.6	55.3	62.6
55.5	60.4	84.2	68.9	65.2	70	56.6	56.3	71.9	60.3
54	65.2	56.3	69.6	68.8	66	63.5	64.9		

We compare the WDEx distribution with some well-known competing distributions including the beta exponential (BEx) [25], gamma (G), exponentiated Weibull (EW) [26], Harris extend exponential (HEEx) [27], alpha power exponentiated-exponential (APEEx) [28], Frechet–Weibull mixture exponential (FWMEx) [29], Marshall–Olkin alpha-power exponential (MOAPEx) [30], Marshall Olkin exponential (MOEx) [31], Frechet–Weibull (FW) [32], Weibull (W), alpha-power exponential (APEx) [33], transmuted exponential (TEx) [34], EEx, and Ex distributions. The competing models are compared by using some discrimination measures such as Anderson–Darling (AD), Cramér–von Mises (CM), and Kolmogorov–Smirnov (KS) with its  $p$ -value (KS  $p$ -value).

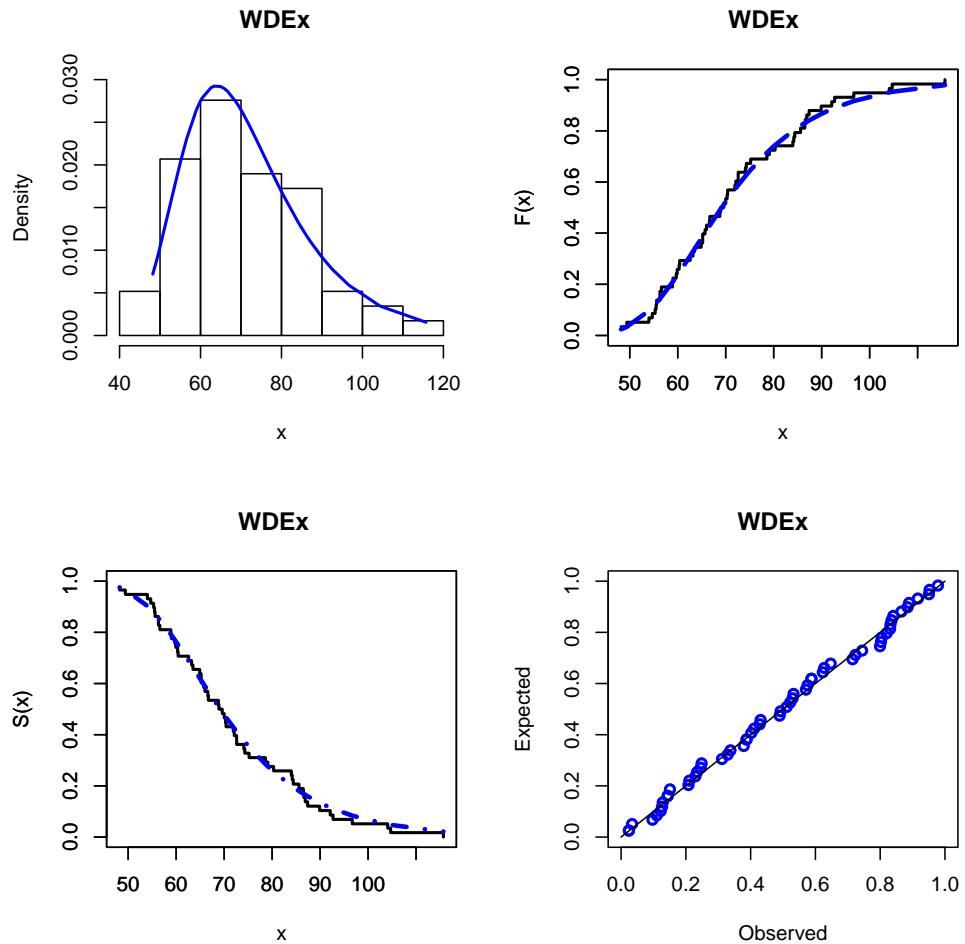
The MLEs and the statistical metrics are calculated using the R software©. Table 6 provides analytical metrics along with the MLEs and their standard errors (SEs). The findings in Table 6 suggest that the WDEx offers better fits than other competitive models and it can be selected as an appropriate model to evaluate the analyzed insurance data.

The fitted PDF, CDF, SF, and probability-probability (P-P) graphs of the WDEx model are displayed in Figure 6, whereas the total time on test (TTT) and WDEx HRF plots for the analyzed data are depicted in Figure 7. The TTT plots confirms that the data has an increasing HRF. We conclude that the WDEx distribution is suitable for modeling the insurance data.

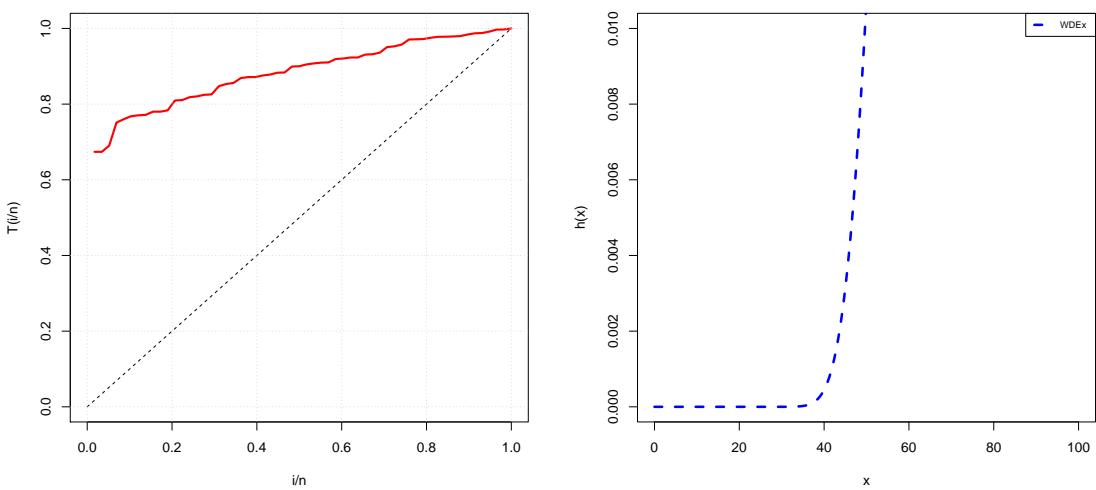
Moreover, the nine estimation approaches mentioned before in Section 5 are adopted to estimate the WDEx parameters. Table 7 records the calculations of the WDEx parameters using such estimation techniques and the empirical values for insurance data. The P-P graphs and histogram of insurance data with the fitted WDEx density for different estimation techniques are shown in Figure 8. Figure 9 shows the histogram of the data with the fitted WDEx PDF for different estimators.

**Table 6.** The analytical measures and the MLEs of the WDE<sub>x</sub> distribution and other competing models.

Model	AD	CM	KS	KS <i>p</i> -value	Estimates (SEs)
WDE <sub>x</sub>	0.22888	0.03088	0.05815	0.98949	$\hat{\alpha} = 132.362$ (56.1205) $\hat{\lambda} = 0.07249$ (0.00691)
FW	0.38156	0.04884	0.06949	0.94209	$\hat{\alpha} = 4.46474$ (66.2443) $\hat{\beta} = 3.17550$ (115.269) $\hat{\lambda} = 25.3428$ (579.785) $\hat{k} = 1.25626$ (18.6392)
BEx	0.27609	0.04162	0.07932	0.85886	$\hat{\alpha} = 75.9965$ (84.8089) $\hat{b} = 1.91398$ (1.80320) $\hat{\lambda} = 0.05573$ (0.03166)
EEx	0.23791	0.03101	0.08120	0.83898	$\hat{\alpha} = 232.474$ (126.888) $\hat{\lambda} = 0.08431$ (0.00884)
G	0.45378	0.07394	0.08162	0.83451	$\hat{\alpha} = 24.6226$ (4.54042) $\hat{b} = 0.34427$ (0.06413)
EW	0.23348	0.03278	0.08179	0.83263	$\hat{\beta} = 1.15364$ (0.11836) $\hat{\lambda} = 0.03802$ (0.02282) $\hat{\theta} = 106.742$ (69.8959)
HEEx	0.31048	0.04283	0.08205	0.82975	$\hat{\alpha} = 3.80452$ (0.75598) $\hat{\lambda} = 0.00662$ (0.00113) $\hat{\theta} = 2115870$ (11864.4)
APEEx	0.22899	0.03147	0.08234	0.82653	$\hat{\alpha} = 1.22817$ (1.95307) $\hat{\alpha} = 0.08631$ (0.01593) $\hat{c} = 248.897$ (168.117)
FWMEx	0.43428	0.06586	0.09245	0.70449	$\hat{\alpha} = 0.01001$ (0.01317) $\hat{\alpha} = 18.1993$ (114.082) $\hat{\lambda} = 40.6229$ (6.75603) $\hat{k} = 0.47024$ (2.94764)
MOAPEx	0.77269	0.12783	0.09341	0.69217	$\hat{\alpha} = 413.107$ (876.070) $\hat{\lambda} = 0.11824$ (0.01269) $\hat{\theta} = 674.591$ (637.159)
MOEx	0.77594	0.12836	0.09397	0.68499	$\hat{\alpha} = 4174.02$ (3747.42) $\hat{\lambda} = 0.11858$ (0.01261)
W	1.19149	0.18595	0.12496	0.32540	$\hat{\alpha} = 4.88793$ (0.46156) $\hat{b} = 77.6595$ (2.21593)
APEx	0.42964	0.06957	0.28502	0.00016	$\hat{\alpha} = 1432.64$ (802.869) $\hat{\lambda} = 0.00352$ (0.00233)
TEX	0.43477	0.07062	0.39111	< 0.00001	$\hat{\lambda} = -1.00000$ (0.51995) $\hat{\theta} = 0.02028$ (0.00262)
Ex	0.45213	0.07366	0.49029	< 0.00001	$\hat{\lambda} = 0.01398$ (0.00183)



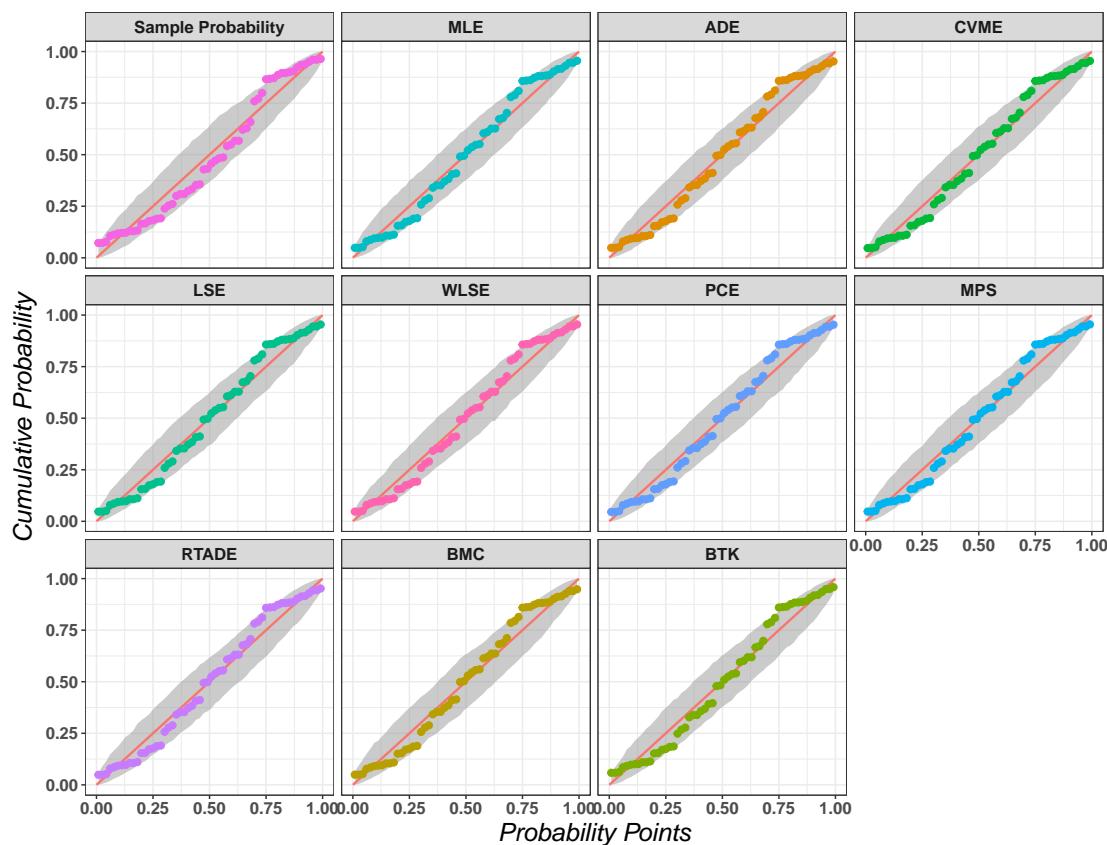
**Figure 6.** The fitted WDEEx PDF, CDF, SF, and P-P plots for insurance data.



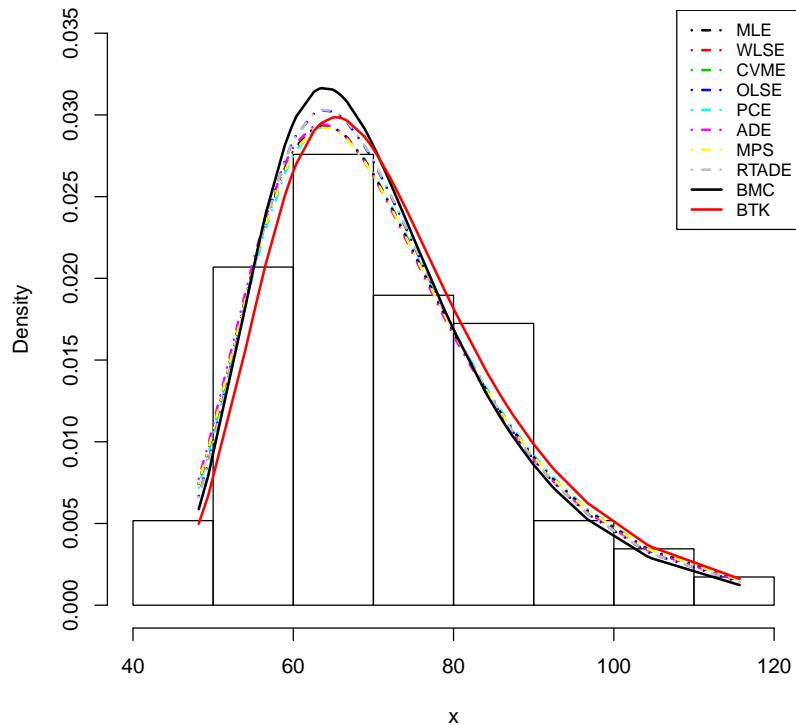
**Figure 7.** The HRF plot of the WDEEx model and TTT plot for insurance data set.

**Table 7.** Different estimates of the WDE<sub>x</sub> parameters using several estimators along with goodness-of-fit measures for insurance data.

Method	$\hat{\alpha}$	$\hat{\lambda}$	AD	CM	KS	KS p-value
MLE	132.36153	0.07249	0.22888	0.03088	0.05815	0.98949
WLSE	132.36153	0.07269	0.22855	0.03084	0.06117	0.98173
CVME	132.36173	0.07270	0.22853	0.03083	0.06138	0.98107
OLSE	132.36145	0.07273	0.22849	0.03083	0.06173	0.97992
PCE	132.36153	0.07295	0.22812	0.03078	0.06516	0.96630
ADE	155.92600	0.07513	0.22928	0.03068	0.06857	0.94795
MPS	130.36158	0.07244	0.22857	0.03086	0.06015	0.98469
RTADE	157.11897	0.07521	0.22939	0.03068	0.06846	0.94860
BMC	194.1689	0.07856	0.23413	0.03148	0.08025	0.84907
BTK	160.4825	0.07402	0.49763	0.07956	0.08157	0.83500



**Figure 8.** The P-P plots of the WDE<sub>x</sub> model using different estimators for insurance data.



**Figure 9.** The histogram of insurance data with the fitted WDEEx densities using different estimators.

## 8. Conclusions

In this paper, we studied a new extension of the exponential model that exhibits upside-down bathtub, decreasing, decreasing-constant, increasing, and increasing-constant hazard rates. The studied model is called the weighted-exponential (WDEEx) distribution. Some basic distributional characteristics are derived. Five risk measures are calculated for the WDEEx distribution. Eight classical estimation methods along with the Bayesian estimation are utilized to estimate the WDEEx parameters. The simulation results show that the proposed estimators provide good performance. Moreover, the practical applicability of the WDEEx distribution is addressed using a real-life insurance data set, showing its superiority fit over fourteen competing exponential models.

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## Conflict of interest

The authors declare there is no conflict of interest.

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