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Theory article

Blowup for regular solutions and C^1 solutions of the two-phase model in \mathbb{R}^N with a free boundary

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Abstract: In this paper, under the assumption of an initial bounded region $\Omega(0)$, we establish the blowup phenomenon of the regular solutions and C^1 solutions to the two-phase model in \mathbb{R}^N . If the total energy *E* and the total mass M > 0 satisfy

$$\max_{\vec{x}_0 \in \partial \Omega(0)} \sum_{i=1}^{N} u_i^2(0, \vec{x}_0) < \frac{\min\{2, N(\Gamma - 1), N(\gamma - 1)\}E}{M},$$

where $E = \int_{\Omega(0)} \left(\frac{1}{2}n |\vec{u}|^2 + \frac{1}{2}\rho |\vec{u}|^2 + \frac{1}{\Gamma-1}n^{\Gamma} + \frac{1}{\gamma-1}\rho^{\gamma} \right) dV$ and $M = \int_{\Omega(0)} (n+\rho)dV > 0$, then the blowup of the solutions to the two-phase model will be formed in finite time in \mathbb{R}^N . Furthermore, under the assumptions that the radially symmetric initial data and initial density contain vacuum states, the blowup of the smooth solutions to the two-phase model will be formed in finite time in \mathbb{R}^N ($N \ge 2$).

Keywords: two-phase model; singularity formation; free boundary problems; regular solutions; vacuum

Mathematics Subject Classification: 35B44, 35R35, 76T10

1. Introduction

The compressible inviscid liquid-gas two-phase model in \mathbb{R}^N that will be considered is as follows:

$$\begin{cases} n_t + div(n\vec{u}) = 0, \\ \rho_t + div(\rho\vec{u}) = 0, \\ [(\rho + n)\vec{u}]_t + div[(\rho + n)\vec{u} \otimes \vec{u}] + \nabla P(n, \rho) = 0, \end{cases}$$
(1.1)

where $\vec{u} = \vec{u}(t, \vec{x})$, $P = P(n, \rho) = n^{\Gamma} + \rho^{\gamma}$ are the velocity and pressure while $n = n(t, \vec{x}) \ge 0$, $\rho = \rho(t, \vec{x}) \ge 0$ are the densities of two phases, where $\gamma, \Gamma > 1$.

Euler equations have been used as one of the basic models for studying fluids, plasmas, atmospheric dynamics, and condensed matter in [2, 5–7, 11, 30], and the blowup analysis for these equations is carried out in [3, 4, 12, 17, 21, 25, 26, 28, 34, 36]. In 1985, Sideris [25] constructed the functional

$$F(0) = \int_{\mathbb{R}^3} \vec{x} \cdot \rho \vec{u} d\vec{x}$$
(1.2)

to prove that the C^1 solutions of the three-dimensional compressible Euler equations will blow up in a finite time when the initial functional F(0) is sufficiently large. Yuen [35] used the energy method to consider the blowup results of the C^1 solutions and the weakened regular solutions of the Euler equations in \mathbb{R}^N . In [18], the authors rewrote the system in the form of a quasilinear wave equation about the density ρ to study the blowup of solutions to Euler equations. Liu, Wang and Yuen studied the blowup results of solutions to the compressible Euler equations with time-dependent damping with vacuum and C^1 solutions of the irrotational compressible Euler equations with time-dependent damping in [19]. In [24], the author studied the three-dimensional Euler equations with a free boundary subjected to tension.

In the two-phase fluid, because of its wide application in aerospace, micro-technology, chemical engineering and other fields, it has aroused many researchers' interest. Zuber studied the two-phase model firstly in [39]. The motion of liquid and gas mixture is studied by the two-phase model in [16]. In [23], the authors considered a hyperbolic two-phase model. For existence, asymptotic and uniqueness of global weak solutions to the two-phase flow model with vacuum, Yao, Zhang and Zhu used the line method and a priori estimate to obtain relevant results in [31-33], and there are also studies of these issues in [1, 8, 13–15, 27]. In [37], under the assumption of H^2 -norm of the initial perturbation with a constant state is sufficiently small and L^1 -norm is bounded, Zhang and Zhu studied the global existence of Cauchy problem to viscous liquid-gas two-phase flow in three dimensions. Furthermore, for the Cauchy problem of 3D inviscid liquid-gas two-phase flow, Zhang considered the optimal $L^P - L^2$ $(1 \le p < \frac{6}{5})$ time decay rates of the solutions with the damping on the qualitative behaviors in [38]. Wen and Zhu [29] considered global existence of weak solutions to two-fluid about the Dirichlet problem in one dimension. In [10], Dong et al. considered the energy integration method to prove the singularity of the smooth solutions to the Cauchy problem for the viscous two-phase model in arbitrary dimensions. Furthermore, Dong and Yuen provided the blowup phenomena of self-similar solutions for the inviscid liquis-gas two-phase flow [9]. By introducing the definition of regular solutions, Makino and Perthame studied the blowup phenomena of radical symmetric solutions to the Euler-Poisson equations with compact support and with a repulsive force and an attractive force in [20, 22].

2. Materials and methods

In this paper, we study the blowup results of regular solutions and C^1 solutions of the two-phase model in \mathbb{R}^N with a free boundary. Our method depends on the energy integration method and a quasi-

linear wave equation about ρ , the singularity of the two-phase model will form in finite time.

3. Results

3.1. Main theorems

In this section, we give the blowup results of regular solutions and C^1 solutions of the two-phase model in \mathbb{R}^N with a free boundary. Firstly, we introduce the definition of regular solutions to the two-phase (1.1).

Definiton 3.1. (Weakened regular solution) A solution (n, ρ, \vec{u}) to the system (1.1) which is in \mathbb{R}^N and in the non-vacuum region is regular for $1 < \Gamma < 2$ or $1 < \gamma < 2$, if

$$\left(n^{\frac{\Gamma-1}{\eta}}, \rho^{\frac{\gamma-1}{\eta}}, \vec{u}\right) \in C^1,$$
(3.1)

with a fixed constant $\eta > 1$.

In the following, we state the blowup results of the two-phase model with a free boundary in \mathbb{R}^N .

Theorem 3.1. Assume that the fluid enters a bounded open region $\Omega(t) \subseteq \mathbb{R}^N$, with the contacting vacuum boundary $\partial \Omega(t)$. Suppose that the (n, ρ, \vec{u}) is a weakened regular solution on $[0, T) \times \Omega(t)$ of the two-phase model (1.1) in \mathbb{R}^N . If

$$\max_{\vec{x}_0 \in \partial \Omega(0)} \sum_{i=1}^{N} u_i^2(0, \vec{x}_0) < \frac{\min\{2, N(\Gamma - 1), N(\gamma - 1)\}E}{M},$$
(3.2)

where the total energy $E = \int_{\Omega(0)} \left(\frac{1}{2}n \left| \vec{u} \right|^2 + \frac{1}{2}\rho \left| \vec{u} \right|^2 + \frac{1}{\Gamma-1}n^{\Gamma} + \frac{1}{\gamma-1}\rho^{\gamma} \right) dV$ and the total mass $M = \int_{\Omega(0)} (n+\rho) dV > 0$, then the solutions of the two-phase model will blow up in finite time T.

Furthermore, we consider the two-phase model (1.1) with the pressure $P = P(n, \rho) = n^{\gamma} + \rho^{\gamma}$. In this paper, we will rewrite the two-phase model (1.1) by forming a quasi-linear wave equation about the density ρ and studying the blowup results for solutions of the two-phase model in $\mathbb{R}^N (N \ge 2)$ with the initial data

$$t = 0: n = n_0(\vec{x}), \ \rho = \rho_0(\vec{x}), \ \vec{u} = \vec{u}_0(\vec{x}).$$
 (3.3)

Introducing radial symmetry, the initial data become

$$n_0(\vec{x}) = n_0(r), \quad \rho_0(\vec{x}) = \rho_0(r), \quad \vec{u}_0(\vec{x}) = \frac{\vec{x}}{r} V_0(r),$$
 (3.4)

where $r = (\sum_{i=1}^{N} x_i^2)^{\frac{1}{2}}$.

In the following, we give the blowup results of the two-phase model in $\mathbb{R}^N (N \ge 3)$.

Theorem 3.2. Assume that the initial density $n_0 \ge 0$ and $\rho_0 \ge 0$ in (3.3), and $(n_0, \rho_0, \vec{u_0}) \in H^3(\mathbb{R}^N)$. Consider the solutions (n, ρ, \vec{u}) of the two-phase model (1.1) in $\mathbb{R}^N (N \ge 3)$. If the initial conditions

$$n_0(0) = 0, \quad \rho_0(0) = 0,$$
 (3.5)

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$$\int_{\mathbb{R}^{N}} (n_0 + \rho_0)(r) dx > 0, \tag{3.6}$$

and

$$-\int_{\mathbb{R}^{N}} \frac{(1+r)(n_{0}+\rho_{0})V_{0}}{r^{2}e^{r}} dx \ge \sqrt{\frac{1}{2^{\gamma-1}(1+\gamma)}} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}}\right]^{\gamma-1} \left(\int_{\mathbb{R}^{N}} \frac{(n_{0}+\rho_{0})}{re^{r}} dx\right)^{\frac{\gamma+1}{2}}$$
(3.7)

are satisfied, where $\frac{\pi^{\frac{N}{2}}}{\Gamma(1+\frac{N}{2})}$ is the volume of the unit ball in \mathbb{R}^N , then the solutions will blow up on or before the finite time $\frac{2H(0)^{-\frac{\gamma-1}{2}}}{(\gamma-1)C_0}$, where $H(0) = \int_{\mathbb{R}^N} \frac{(n_0+\rho_0)}{re^r} dx$ and $C_0 = \sqrt{\frac{1}{2^{\gamma-1}(1+\gamma)} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}}\right]^{\gamma-1}}$.

Then, under the appropriate assumptions, we will give the blowup results of the two-phase model in the two dimensional cases.

Theorem 3.3. Assume that the initial density $n_0 \ge 0$, $\rho_0 \ge 0$ in (3.3), and $(n_0, \rho_0, \vec{u_0}) \in H^3(\mathbb{R}^2)$. Let $K_0(r)$ be the modified Bessel function

$$K_0(r) = \int_0^\infty e^{-rcosht} dt.$$
(3.8)

Consider the solution (n, ρ, \vec{u}) of the two-phase model (1.1) in \mathbb{R}^2 . If the initial conditions

$$n_0(0) = 0, \quad \rho_0(0) = 0,$$
 (3.9)

$$\int_{\mathbb{R}^2} (n_0 + \rho_0)(r) dx > 0, \tag{3.10}$$

and

$$\int_{\mathbb{R}^{2}} (n_{0} + \rho_{0}) V_{0} K_{0}'(r) dx \ge \sqrt{\frac{1}{2^{\gamma - 1}(1 + \gamma)}} \frac{\left(\int_{\mathbb{R}^{2}} (n_{0} + \rho_{0})(r) K_{0}(r) dx\right)^{\frac{\gamma - 1}{2}}}{\left(\int_{\mathbb{R}^{2}} K_{0}(r) dx\right)^{\frac{\gamma - 1}{2}}}$$
(3.11)

are satisfied, then the solutions will blow up on or before the finite time $\frac{2G(0)^{-\frac{\gamma-1}{2}}}{(\gamma-1)C_1}$, where $G(0) = \int_{\mathbb{R}^2} (n_0 + \rho_0) K_0(r) dx$ and $C_1 = \sqrt{\frac{1}{2^{\gamma-1}(\gamma+1)}} \left(\int_{\mathbb{R}^2} K_0(r) dx \right)^{-\frac{\gamma-1}{2}}$.

Remark 3.1. For the two-phase model in \mathbb{R}^N with pressure $P = P(n, \rho) = (n + \rho)^{\gamma}$, we can also obtain the same blowup results.

3.2. Blowup for the two-phase model with a free boundary

In this section, we will show the proof of the blowup result of solutions to the two-phase model in \mathbb{R}^N . Firstly, we give some lemmas for the conserved energy and the properties of the local second inertia function to the two-phase model.

Lemma 3.1. Assume the density n and ρ have compact support in the region $\Omega(t)$ for non-trivial C^1 solutions of the two-phase model (1.1) in \mathbb{R}^N , where $\Omega(t)$ is a moving region, and the region $\Omega(0)$ is bounded by the contacting vacuum boundary, then

$$M(t) = M(0), (3.12)$$

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and

$$0 < E = E(t) = \int_{\Omega(t)} \left(\frac{1}{2} n \left| \vec{u} \right|^2 + \frac{1}{2} \rho \left| \vec{u} \right|^2 + \frac{1}{\Gamma - 1} n^{\Gamma} + \frac{1}{\gamma - 1} \rho^{\gamma} \right) dV < \infty,$$
(3.13)

that is, the total energy energy E(t) is conserved, where $M(t) = \int_{\Omega(t)} (n+\rho) dV$ represents the total mass. *Proof.* By $(1.1)_1$ and $(1.1)_2$, we obtain

$$\frac{d}{dt}M(t) = \int_{\Omega(t)} (n_t + \rho_t)dV = -\int_{\Omega(t)} div(n\vec{u}) + div(\rho\vec{u})dV = 0.$$
(3.14)

Therefore, M(t) = M(0).

Multiplying \vec{u} on both sides of the Eq (1.1)₃ and integrating over $\Omega(t)$, we obtain

$$\int_{\Omega(t)} (n\vec{u})_t \cdot \vec{u} dV + \int_{\Omega(t)} div(n\vec{u} \otimes \vec{u}) \cdot \vec{u} dx + \int_{\Omega(t)} (\rho\vec{u})_t \cdot \vec{u} dV + \int_{\Omega(t)} div(\rho\vec{u} \otimes \vec{u}) \cdot \vec{u} dV + \int_{\Omega(t)} \nabla P(n,\rho) \cdot \vec{u} dV = 0.$$
(3.15)

On the one hand, from $(1.1)_1$, we have

$$-\frac{1}{2}n_t \left| \vec{u} \right|^2 - \frac{1}{2} \left| \vec{u} \right|^2 div(n\vec{u}) = 0.$$
(3.16)

Due to

$$\left(\frac{1}{2}n\left|\vec{u}\right|^{2}\right)_{t} = (n\vec{u})_{t} \cdot \vec{u} - \frac{1}{2}n_{t}\left|\vec{u}\right|^{2},$$
(3.17)

we obtain

$$\int_{\Omega(t)} (n\vec{u})_t \cdot \vec{u} dx + \int_{\Omega(t)} div(n\vec{u} \otimes \vec{u}) \cdot \vec{u} dV$$

$$= \int_{\Omega(t)} \left(\frac{1}{2}n\left|\vec{u}\right|^2\right)_t dV + \int_{\Omega(t)} \frac{1}{2}n_t \left|\vec{u}\right|^2 dV + \int_{\Omega(t)} \vec{u} \cdot \left[\nabla \cdot (n\vec{u} \otimes \vec{u})\right] dV$$

$$= \int_{\Omega(t)} \left(\frac{1}{2}n\left|\vec{u}\right|^2\right)_t dV - \int_{\Omega(t)} \frac{1}{2} \left|\vec{u}\right|^2 \nabla \cdot (n\vec{u}) dV + \int_{\Omega(t)} \vec{u} \cdot \left[\nabla \cdot (n\vec{u} \otimes \vec{u})\right] dV. \quad (3.18)$$

We note that

$$\vec{u} \cdot \left[\nabla \cdot (n\vec{u} \otimes \vec{u})\right] = \sum_{i, j=1}^{N} u_i \left[\partial_j \left(\rho u_j\right) u_i + n u_j \partial_j u_i\right] = \left|\vec{u}\right|^2 \sum_{i=1}^{N} \partial_i \left(\rho u_i\right) + \sum_{i, j=1}^{N} n u_i u_j \partial_j u_i.$$
(3.19)

Thus, we get

$$\begin{split} &\int_{\Omega(t)} \left[-\frac{1}{2} \left| \vec{u} \right|^2 \nabla \cdot (n\vec{u}) + \vec{u} \cdot \left[\nabla \cdot (n\vec{u} \otimes \vec{u}) \right] \right] dV \\ &= \int_{\Omega(t)} \left[-\frac{1}{2} \left| \vec{u} \right|^2 \sum_{i=1}^N \partial_i \left(nu_i \right) + \left| \vec{u} \right|^2 \sum_{i=1}^N \partial_i \left(nu_i \right) \sum_{i, j=1}^N nu_i u_j \partial_j u_i \right] dV \end{split}$$

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$$= \int_{\Omega(t)} \left[\frac{1}{2} \left| \vec{u} \right|^2 \sum_{i=1}^N \partial_i (nu_i) + \sum_{i, j=1}^N nu_i u_j \partial_j u_i \right] dV$$

$$= \int_{\Omega(t)} \left[-\sum_{i=1}^N \vec{u} \cdot \partial_i \vec{u} nu_i + \sum_{i, j=1}^N nu_i u_j \partial_j u_i \right] dV$$

$$= \int_{\Omega(t)} \left[-\sum_{i, j=1}^N nu_i u_j \partial_i u_j + \sum_{i, j=1}^N nu_i u_j \partial_j u_i \right] dV$$

$$= 0.$$
(3.20)

Therefore, we obtain

$$\int_{\Omega(t)} (n\vec{u})_t \cdot \vec{u} dV + \int_{\Omega(t)} div(n\vec{u} \otimes \vec{u}) \cdot \vec{u} dV = \int_{\Omega(t)} \left(\frac{1}{2}n\left|\vec{u}\right|^2\right)_t dV.$$
(3.21)

Similarly, we get

$$\int_{\Omega(t)} (\rho \vec{u})_t \cdot \vec{u} dV + \int_{\Omega(t)} div(\rho \vec{u} \otimes \vec{u}) \cdot \vec{u} dV = \int_{\Omega(t)} \left(\frac{1}{2}\rho \left|\vec{u}\right|^2\right)_t dV.$$
(3.22)

On the other hand,

$$P_{t} = \Gamma n^{\Gamma-1} \partial_{t} n + \gamma \rho^{\gamma-1} \partial_{t} \rho$$

$$= \Gamma n^{\Gamma-1} \left[-\sum_{i=1}^{N} \partial_{i} (nu_{i}) \right] + \gamma \rho^{\gamma-1} \left[-\sum_{i=1}^{N} \partial_{i} (\rho u_{i}) \right]$$

$$= -\sum_{i=1}^{N} \partial_{i} (n^{\Gamma}) u_{i} - \sum_{i=1}^{N} \Gamma n^{\Gamma} \partial_{i} u_{i} - \sum_{i=1}^{N} \partial_{i} (\rho^{\gamma}) u_{i} - \sum_{i=1}^{N} \gamma \rho^{\gamma} \partial_{i} u_{i}.$$
(3.23)

Integrating over the region $\Omega(t)$ and applying the integration by parts, we obtain

$$\int_{\Omega(t)} P_t dV = -\int_{\Omega(t)} \sum_{i=1}^N \partial_i (n^{\Gamma}) u_i dV - \int_{\Omega(t)} \sum_{i=1}^N \Gamma n^{\Gamma} \partial_i u_i dV - \int_{\Omega(t)} \sum_{i=1}^N \partial_i (\rho^{\gamma}) u_i dV - \int_{\Omega(t)} \sum_{i=1}^N \gamma \rho^{\gamma} \partial_i u_i dV$$
$$= -\int_{\Omega(t)} \sum_{i=1}^N \partial_i (n^{\Gamma}) u_i dV + \int_{\Omega(t)} \sum_{i=1}^N \Gamma \partial_i (n^{\Gamma}) u_i dV - \int_{\Omega(t)} \sum_{i=1}^N \partial_i (\rho^{\gamma}) u_i dV - \int_{\Omega(t)} \sum_{i=1}^N \gamma \partial_i (\rho^{\gamma}) u_i dV$$
$$= \int_{\Omega(t)} \sum_{i=1}^N (\Gamma - 1) \partial_i (n^{\Gamma}) u_i dV + \int_{\Omega(t)} \sum_{i=1}^N (\gamma - 1) \partial_i (\rho^{\gamma}) u_i dV.$$
(3.24)

Therefore, we have

$$\int_{\Omega(t)} \vec{u} \cdot \nabla P(n,\rho) dV = \frac{1}{\Gamma - 1} \int_{\Omega(t)} \partial_t(n^{\Gamma}) dV + \frac{1}{\gamma - 1} \int_{\Omega(t)} \partial_t(\rho^{\gamma}) dV.$$
(3.25)

Thus, by (3.15), we obtain

$$\frac{d}{dt}E(t) = \frac{d}{dt}\left(\int_{\Omega(t)} \frac{1}{2}n \left|\vec{u}\right|^2 + \frac{1}{2}\rho \left|\vec{u}\right|^2 + \frac{1}{\Gamma - 1}n^{\Gamma} + \frac{1}{\gamma - 1}\rho^{\gamma}dV\right)$$

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$$= \int_{\Omega(t)} (n\vec{u})_t \cdot \vec{u}dV + \int_{\Omega(t)} div(n\vec{u} \otimes \vec{u}) \cdot \vec{u}dx + \int_{\Omega(t)} (\rho\vec{u})_t \cdot \vec{u}dV + \int_{\Omega(t)} div(\rho\vec{u} \otimes \vec{u}) \cdot \vec{u}dV + \int_{\Omega(t)} \nabla P(n,\rho) \cdot \vec{u}dV = 0.$$
(3.26)

Therefore, for non-trivial C^1 solutions, we obtain

$$E(t) = E(0) < \infty. \tag{3.27}$$

The proof is complete.

Next, we show the second derivative of the local second inertia function for the two-phase model (1.1) in \mathbb{R}^N for the solutions on $[0, T) \times \Omega(t)$.

Before stating the following lemma, we first give some physical quantities as follows:

$$F(t) = \int_{\Omega(t)} n\vec{u} \cdot xdV + \int_{\Omega(t)} \rho\vec{u} \cdot xdV = F_n(t) + F_\rho(t), \qquad (3.28)$$

$$H(t) = \frac{1}{2} \int_{\Omega(t)} n \left| \vec{x} \right|^2 dV + \frac{1}{2} \int_{\Omega(t)} \rho \left| \vec{x} \right|^2 dV = H_n(t) + H_\rho(t),$$
(3.29)

and

$$\begin{split} E(t) &= \int_{\Omega(t)} \frac{1}{2} n \left| \vec{u} \right|^2 dV + \int_{\Omega(t)} \frac{1}{2} \rho \left| \vec{u} \right|^2 dV + \frac{1}{\Gamma - 1} \int_{\Omega(t)} n^{\Gamma} dV + \frac{1}{\gamma - 1} \int_{\Omega(t)} \rho^{\gamma} dV \\ &= E_{kn}(t) + E_{k\rho}(t) + E_{in}(t) + E_{i\rho}(t), \end{split}$$
(3.30)

where $F_n(t)$ and $F_\rho(t)$, $H_n(t)$ and $H_\rho(t)$, $E_{kn}(t)$ and $E_{k\rho}(t)$, $E_{in}(t)$ and $E_{i\rho}(t)$ represent the momentum weight, the local second inertia, the kinetic energy and the internal energy for the two-phase fluid, respectively.

Lemma 3.2. For the two-phase model (1.1) in \mathbb{R}^N , we have

$$H'(t) = F(t),$$
 (3.31)

and

$$H^{''}(t) = 2E_{kn}(t) + 2E_{k\rho}(t) + N(\Gamma - 1)E_{in}(t) + N(\gamma - 1)E_{i\rho}(t).$$
(3.32)

Proof. Differentiating H(t) with respect to t, we obtain

$$H'(t) = \frac{1}{2} \int_{\Omega(t)} n_t \left| \vec{x} \right|^2 dV + \frac{1}{2} \int_{\Omega(t)} \rho_t \left| \vec{x} \right|^2 dV = H'_n(t) + H'_\rho(t).$$
(3.33)

Applying the integration by parts to (3.33) and using $(1.1)_1$, we get

$$H'_{n}(t) = -\frac{1}{2} \int_{\Omega(t)} div(n\vec{u}) \left|\vec{x}\right|^{2} dV = -\frac{1}{2} \int_{\Omega(t)} \sum_{i,j}^{N} \partial_{i}(nu_{i}) x_{j}^{2} dV$$

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$$= \frac{1}{2} \int_{\Omega(t)} \sum_{i,j}^{N} 2nu_i x_j dV = \int_{\Omega(t)} \sum_{i}^{N} nu_i x_i dV$$
$$= \int_{\Omega(t)} \vec{x} \cdot n\vec{u} dV = F_n(t).$$
(3.34)

Similarly, we have

$$H'_{\rho}(t) = F_{\rho}(t).$$
 (3.35)

Therefore, we obtain

$$H'(t) = F(t).$$
 (3.36)

Next, we calculate H''(t) as follows:

$$H''(t) = F'(t) = \int_{\Omega(t)} (n\vec{u})_t \cdot \vec{x} dV + \int_{\Omega(t)} (\rho\vec{u})_t \cdot \vec{x} dV$$

$$= -\int_{\Omega(t)} div(n\vec{u} \otimes \vec{u}) \cdot \vec{x} dV - \int_{\Omega(t)} div(\rho\vec{u} \otimes \vec{u}) \cdot \vec{x} dV - \int_{\Omega(t)} \nabla P(n,\rho) \cdot \vec{x} dV. \quad (3.37)$$

By using the integration by parts, we get

$$\int_{\Omega(t)} div(n\vec{u} \otimes \vec{u}) \cdot \vec{x} dV = \int_{\Omega(t)} \sum_{i,j}^{N} \partial_i (nu_i u_j) x_j dV = -\int_{\Omega(t)} \sum_{i,j}^{N} nu_i u_j \partial_i x_j dV$$
$$= -\int_{\Omega(t)} n\vec{u} \cdot \vec{u} dV = -\int_{\Omega(t)} n \left|\vec{u}\right|^2 dV.$$
(3.38)

Similarly, we have

$$\int_{\Omega(t)} div(\rho \vec{u} \otimes \vec{u}) \cdot \vec{x} dV = -\int_{\Omega(t)} \rho \left| \vec{u} \right|^2 dV,$$
(3.39)

and

$$\int_{\Omega(t)} \nabla P(n,\rho) \cdot \vec{x} dV = -N \int_{\Omega(t)} n^{\Gamma} dV - N \int_{\Omega(t)} \rho^{\gamma} dV.$$
(3.40)

From (3.38) to (3.40) and Gauss formula, we obtain

$$H''(t) = F'(t) = \int_{\Omega(t)} n \left| \vec{u} \right|^2 dV + \int_{\Omega(t)} \rho \left| \vec{u} \right|^2 dV + N \int_{\Omega(t)} n^{\Gamma} dV + N \int_{\Omega(t)} \rho^{\gamma} dV$$

= $2E_{kn}(t) + 2E_{k\rho}(t) + N(\Gamma - 1)E_{in}(t) + N(\gamma - 1)E_{i\rho}(t).$ (3.41)

The proof is complete.

Using the above lemmas, we will give the proof of the blowup results of the solutions to the twophase model with a free boundary in \mathbb{R}^N .

Proof of Theorem 3.1. Introducing

$$\varphi = n^{\frac{\Gamma-1}{\eta}}, \omega = \rho^{\frac{\gamma-1}{\eta}}, \tag{3.42}$$

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for $n(t, x) \neq 0$ and $\rho(t, x) \neq 0$, we transform the momentum equations (1.1) into

$$(\varphi^{\frac{\eta}{\Gamma-1}} + \omega^{\frac{\eta}{\gamma-1}}) \left(\frac{\partial u_i}{\partial t} + \sum_{l=1}^N u_l \frac{\partial u_i}{\partial x_l} \right) + \frac{\partial}{\partial x_i} (\varphi^{\frac{\eta\Gamma}{\Gamma-1}} + \omega^{\frac{\eta\gamma}{\gamma-1}}) = 0.$$
(3.43)

Simplifying the above equation, we have

$$\left(\varphi^{\frac{\eta}{\Gamma-1}} + \omega^{\frac{\eta}{\gamma-1}}\right) \left(\frac{\partial u_i}{\partial t} + \sum_{l=1}^N u_l \frac{\partial u_i}{\partial x_l}\right) + \frac{\eta\Gamma}{\Gamma-1} \varphi^{\frac{\eta\Gamma}{\Gamma-1}-1} \frac{\partial}{\partial x_i} \varphi + \frac{\eta\gamma}{\gamma-1} \omega^{\frac{\eta\gamma}{\gamma-1}-1} \frac{\partial}{\partial x_i} \omega = 0, \quad (3.44)$$

where $\eta > 1$ is an arbitrary constant, $(\eta - 1)\Gamma + 1 - \eta > 0$ and $(\eta - 1)\gamma + 1 - \eta > 0$.

We study the solutions near the contacting vacuum boundary point $\vec{x}_0(t)$ in the region $\Omega(t)$. Firstly, we consider the local second inertial function

$$H(t) = \frac{1}{2} \int_{\Omega(t)} n \left| \vec{x} \right|^2 dV + \frac{1}{2} \int_{\Omega(t)} \rho \left| \vec{x} \right|^2 dV.$$
(3.45)

By Lemma 3.2, we obtain

$$H''(t) = 2E_{kn}(t) + 2E_{k\rho}(t) + N(\Gamma - 1)E_{in}(t) + N(\gamma - 1)E_{i\rho}(t)$$

$$\geq \min\{2, N(\Gamma - 1), N(\gamma - 1)\}E(t)$$

$$= C_1E,$$
(3.46)

with the conserved total energy $E = \int_{\Omega(0)} \left(\frac{1}{2}n \left| \vec{u} \right|^2 + \frac{1}{2}\rho \left| \vec{u} \right|^2 + \frac{1}{\Gamma-1}n^{\Gamma} + \frac{1}{\gamma-1}\rho^{\gamma} \right) dV$ by Lemma 3.1, where $C_1 = \min\{2, N(\Gamma-1), N(\gamma-1)\}.$

Therefore, we obtain

$$H(t) \ge H(0) + \dot{H}(0)t + \frac{C_1 E}{2}t^2.$$
(3.47)

For an arbitrary point $\vec{x}_0(t) \in \partial \Omega(t)$ with the contacting vacuum, by applying Lemma 3.1 and $\rho(0, \vec{x}_0) = 0$, we obtain

$$\frac{\partial u_i}{\partial t} + \sum_{l=1}^N u_l \frac{\partial u_l}{\partial x_l} = 0.$$
(3.48)

We consider the governing differential dynamic system at the contacting vacuum point $\vec{x}_0(t)$:

$$\begin{cases} \frac{D^2 x_{0,i}(t)}{Dt^2} = 0, & \text{for } i = 1, 2, \cdots, N, \\ x_{0,i}(0, \vec{x}_0) = x_{0,i}, & \dot{x}_{0,i}(0, \vec{x}_0) = u_i(0, \vec{x}_0), \end{cases}$$
(3.49)

where $\frac{dx_{0,i}(t)}{dt} = u_i(t, \vec{x}_0(t)).$

The solutions of the above system are as follows:

$$x_{0,i}(t) = x_{0,i} + u_i(0, \vec{x}_0)t.$$
(3.50)

Then, we get

$$u_i(t, \vec{x}_0(t)) = \frac{dx_{0,i}(t)}{dt} = u_i(0, \vec{x}_0).$$
(3.51)

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By applying the Euclidean norm, we obtain

$$\max_{\vec{x}\in\partial\Omega(t)} \left(x_1^2 + x_2^2 + \dots + x_N^2 \right)^{\frac{1}{2}} \ge \max_{\vec{x}\in\Omega(t)} \left(x_1^2 + x_2^2 + \dots + x_N^2 \right)^{\frac{1}{2}} \ge 0.$$
(3.52)

Therefore, we have

$$\max_{\vec{x}_0 \in \partial \Omega(t)} \left| \overrightarrow{x_0} \right|^2 \ge \max_{\vec{x} \in \Omega(t)} \left| \overrightarrow{x} \right|^2.$$
(3.53)

Furthermore, we obtain

$$\frac{1}{2} \max_{\vec{x}_0 \in \partial \Omega(t)} \sum_{i=1}^{N} x_{0,i}^2(t) \int_{\Omega(t)} n dV \ge \frac{1}{2} \int_{\Omega(t)} n \left| \vec{x} \right|^2 dV.$$
(3.54)

Similarly, we obtain

$$\frac{1}{2} \max_{\vec{x}_0 \in \partial \Omega(t)} \sum_{i=1}^{N} x_{0,i}^2(t) \int_{\Omega(t)} \rho dV \ge \frac{1}{2} \int_{\Omega(t)} \rho \left| \vec{x} \right|^2 dV.$$
(3.55)

Therefore,

$$\frac{1}{2} \max_{\vec{x}_0 \in \partial \Omega(t)} \sum_{i=1}^{N} x_{0,i}^2(t) \left(\int_{\Omega(t)} n dV + \int_{\Omega(t)} \rho dV \right) \geq \frac{1}{2} \left(\int_{\Omega(t)} n \left| \vec{x} \right|^2 dV + \int_{\Omega(t)} \rho \left| \vec{x} \right|^2 dV \right) \\
= H(t) \geq H(0) + \dot{H}(0)t + \frac{C_1 E}{2} t^2, \quad (3.56)$$

that is

$$\max_{\vec{x}_0 \in \partial \Omega(t)} \sum_{i=1}^N x_{0,i}^2(t) \ge \frac{2}{M} \left(H(0) + \dot{H}(0)t + \frac{C_1 E}{2} t^2 \right),$$
(3.57)

with the conserved total mass $M = \int_{\Omega(0)} (n + \rho) dV > 0$ in Lemma 3.1.

By (3.50), we have

$$\frac{2}{M} \left(H(0) + \dot{H}(0)t + \frac{C_1 E}{2} t^2 \right) \le \max_{\vec{x}_0 \in \partial \Omega(0)} \sum_{i=1}^N \left(x_{0,i}^2 + 2x_{0,i} \cdot u_i(0, \vec{x}_0)t + u_i^2(0, \vec{x}_0)t^2 \right).$$
(3.58)

When $t \to +\infty$, we obtain

$$\lim_{t \to +\infty} \frac{\frac{2H(0)}{M} + \frac{2\dot{H}(0)}{M}t + \frac{C_1E}{M}t^2}{t^2} \le \lim_{t \to +\infty} \frac{\max_{\vec{x}_0 \in \partial \Omega(0)} \sum_{i=1}^{N} \left(x_{0,i}^2 + 2x_{0,i} \cdot u_i(0, \vec{x}_0)t + u_i^2(0, \vec{x}_0)t^2\right)}{t^2}.$$
 (3.59)

Therefore, we have

$$\frac{C_1 E}{M} \le \max_{\vec{x}_0 \in \partial \Omega(0)} u_i^2(0, \vec{x}_0),$$
(3.60)

that is

$$\frac{\min(2, N(\Gamma - 1), N(\gamma - 1))E}{M} \le \max_{\vec{x}_0 \in \partial \Omega(0)} u_i^2(0, \vec{x}_0),$$
(3.61)

which contradicts with inequality (3.2) in Theorem 3.1. Therefore, we conclude that the solutions for the two-phase model (1.1) in \mathbb{R}^N will blow up in finite time.

This completes the proof.

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3.3. Blowup for the two-phase model with vacuum

For proofing the blowup results of solutions to the two-phase model in $\mathbb{R}^N (N \ge 2)$, we first introduce the properties of the two-phase model.

Lemma 3.3. (*Theorem 2.1 in [18]*) Assume that $n_0 \ge 0$, $\rho_0 \ge 0$ and $(n_0, \rho_0, \vec{u_0}) \in H^3(\mathbb{R}^N)$, then there exists a unique solution (n, ρ, \vec{u}) to the two-phase model on some time interval [0, T), which satisfies

$$n, \rho \in C\left([0, T] \times \mathbb{R}^N\right),\tag{3.62}$$

and

$$\vec{u} \in C([0,T), H^{3}(\mathbb{R}^{N})) \cap C^{1}([0,T), H^{2}(\mathbb{R}^{N})) \cap C^{2}([0,T), H^{1}(\mathbb{R}^{N})).$$
(3.63)

Lemma 3.4. (Lemma 4 in [35]) From the mass equations $(1.1)_1$ and $(1.1)_2$, we obtain

$$n(t, \vec{x}(t)) = n(0, \vec{x}_0) \exp\left(-\int_0^t \nabla \cdot \vec{u} ds\right)$$
(3.64)

and

$$\rho(t, \vec{x}(t)) = \rho(0, \vec{x}_0) \exp\left(-\int_0^t \nabla \cdot \vec{u} ds\right).$$
(3.65)

Next, we give the lemma to show the property of the modified Bessel function $K_0(r)$.

Lemma 3.5. (Lemma 3.1 in [18]) The modified Bessel function $K_0(r) = \int_0^\infty e^{-rcosht} dt$ satisfies

$$\begin{cases} K_0(r) \le \frac{3}{r}, |K'_0(r)| \le \frac{1}{r^2}, & 0 < r < \frac{1}{2}, \\ K_0(r) \le \frac{C_k}{r^k}, |K'_0(r)| \le \frac{C_k}{r^k}, & r > 1, \end{cases}$$
(3.66)

for some constants C_k depending only on k > 1.

Then, we give the proof of the blowup results of solutions to the two-phase model in $\mathbb{R}^N (N \ge 2)$.

Proof of Theorem 3.2. We consider the solution (n, ρ, \vec{u}) that satisfies the conditions in Theorem 3.2. From the two-phase system (1.1), we obtain

$$n_{tt} + \rho_{tt} = -(\nabla \cdot (n\vec{u})_t + \nabla \cdot (\rho\vec{u})_t)$$

= $\Delta P + \nabla \cdot [\nabla \cdot ((n+\rho)\vec{u} \otimes \vec{u})].$ (3.67)

Then, multiplying Eq (3.67) by $\frac{1}{re^r}$ and integrating over \mathbb{R}^N to obtain

$$\frac{d^2}{dt^2} \int_{\mathbb{R}^N} \frac{n+\rho}{re^r} dx = \int_{\mathbb{R}^N} \Delta P \frac{1}{re^r} dx + \int_{\mathbb{R}^N} \nabla \cdot \left[\nabla \cdot \left((n+\rho)\vec{u} \otimes \vec{u}\right)\right] \frac{1}{re^r} dx.$$
(3.68)

By the divergence theorem, we obtain

$$\int_{\mathbb{R}^{N}} \Delta P \frac{1}{re^{r}} dx = \int_{\mathbb{R}^{N}} P \Delta \frac{1}{re^{r}} dx - \lim_{\varepsilon \to 0} \int_{r=\varepsilon} \left(\frac{\partial P}{\partial r} \frac{1}{re^{r}} - \frac{\partial}{\partial r} (\frac{1}{re^{r}}) P \right) ds.$$
(3.69)

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By applying the assumption (3.5) and Lemma 3.4, we obtain $n(t, \vec{0}) \equiv 0, \rho(t, \vec{0}) \equiv 0$. Using the continuity of *P*, we get that for any sufficiently small $\delta > 0$, there is $0 < \varepsilon < 1$, such that $P(t, \vec{x}) < \delta$ while $r = \varepsilon$, then,

$$\int_{r=\varepsilon} \frac{\partial(\frac{1}{re^{r}})}{\partial r} P ds = \int_{r=\varepsilon} \frac{re^{r} + e^{r}}{r^{2}e^{2r}} P ds$$

$$\leq \delta \int_{r=\varepsilon} \frac{re^{r} + e^{r}}{r^{2}e^{2r}} ds = \delta \int_{r=\varepsilon} \frac{(\varepsilon+1)e^{\varepsilon}}{\varepsilon^{2}e^{2\varepsilon}} ds$$

$$= \delta N \alpha(N) \varepsilon^{N-1} \frac{(\varepsilon+1)e^{\varepsilon}}{\varepsilon^{2}e^{2\varepsilon}} = \delta N \alpha(N) \varepsilon^{N-3} \frac{(\varepsilon+1)e^{\varepsilon}}{e^{2\varepsilon}}, \qquad (3.70)$$

where $\alpha(N) = N(N-2)\frac{\pi^{\frac{N}{2}}}{\Gamma(1+\frac{N}{2})}$ and $\frac{\pi^{\frac{N}{2}}}{\Gamma(1+\frac{N}{2})}$ is the volume of the unit ball in \mathbb{R}^N . Therefore, for $N \ge 3$, we obtain

$$\lim_{\varepsilon \to 0} \int_{r=\varepsilon} \frac{\partial(\frac{1}{re^{r}})}{\partial r} P ds = 0.$$
(3.71)

Similarly, we get

$$\lim_{\varepsilon \to 0} \int_{r=\varepsilon} \frac{\partial P}{\partial r} \frac{1}{re^r} ds = 0.$$
(3.72)

Thus, we obtain

$$\int_{\mathbb{R}^N} \Delta P \frac{1}{re^r} dx = \int_{\mathbb{R}^N} P \frac{1}{re^r} dx = \int_{\mathbb{R}^N} \frac{n^{\gamma}}{re^r} dx + \int_{\mathbb{R}^N} \frac{\rho^{\gamma}}{re^r} dx.$$
(3.73)

Using Hölder inequality, we get

$$\left(\int_{\mathbb{R}^{N}} \frac{n}{re^{r}} dx\right)^{\gamma} \le \left(\int_{\mathbb{R}^{N}} \frac{n^{\gamma}}{re^{r}} dx\right) \left(\int_{\mathbb{R}^{N}} \frac{1}{re^{r}} dx\right)^{\gamma-1},\tag{3.74}$$

and

$$\left(\int_{\mathbb{R}^N} \frac{\rho}{re^r} dx\right)^{\gamma} \le \left(\int_{\mathbb{R}^N} \frac{\rho^{\gamma}}{re^r} dx\right) \left(\int_{\mathbb{R}^N} \frac{1}{re^r} dx\right)^{\gamma-1}.$$
(3.75)

Therefore, (3.73) becomes

$$\int_{\mathbb{R}^{N}} \Delta P \frac{1}{re^{r}} dx \geq \frac{1}{(\int_{\mathbb{R}^{N}} \frac{1}{re^{r}} dx)^{\gamma-1}} \left[(\int_{\mathbb{R}^{N}} \frac{n}{re^{r}} dx)^{\gamma} + (\int_{\mathbb{R}^{N}} \frac{\rho}{re^{r}} dx)^{\gamma} \right] \\
\geq \frac{1}{2^{\gamma-1} (\int_{\mathbb{R}^{N}} \frac{1}{re^{r}} dx)^{\gamma-1}} (\int_{\mathbb{R}^{N}} \frac{n+\rho}{re^{r}} dx)^{\gamma} \\
= \frac{1}{2^{\gamma-1}} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}} \right]^{\gamma-1} (\int_{\mathbb{R}^{N}} \frac{n+\rho}{re^{r}} dx)^{\gamma}.$$
(3.76)

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Applying the integration by parts, we obtain

$$\int_{\mathbb{R}^{N}} \nabla \cdot \left[\nabla \cdot \left((n+\rho)\vec{u} \otimes \vec{u}\right)\right] \frac{1}{re^{r}} dx = \int_{\mathbb{R}^{N}} \frac{1}{re^{r}} \sum_{i,j=1}^{n} \frac{\partial^{2} \left((n+\rho)u_{i}u_{j}\right)}{\partial x_{i}\partial x_{j}} dx = \int_{\mathbb{R}^{N}} (n+\rho)V^{2} (\frac{1}{re^{r}})^{''} dx$$
$$= \int_{\mathbb{R}^{N}} (n+\rho)V^{2} [\frac{1}{r} + \frac{2}{r^{2}} + \frac{2}{r^{3}}]e^{-r} dx > 0.$$
(3.77)

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Therefore, we obtain

$$\frac{d^2}{dt^2}H(t) \ge \frac{1}{2^{\gamma-1}} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}} \right]^{\gamma-1} H^{\gamma}(t),$$
(3.78)

where $H(t) = \int_{\mathbb{R}^N} \frac{(n+\rho)(t,r)}{re^r} dx$. Using integration by parts and (3.7), we obtain

$$H'(0) = \frac{d}{dt} \int_{\mathbb{R}^{N}} \frac{(n+\rho)}{re^{r}} dx|_{t=0} = -\int_{\mathbb{R}^{N}} \frac{\nabla \cdot [(n+\rho)\vec{u}]}{re^{r}} dx|_{t=0}$$
$$= -\int_{\mathbb{R}^{N}} \sum_{i=1}^{N} (n+\rho) \frac{x_{i}}{r} V \frac{1+r}{r^{2}e^{r}} \frac{x_{i}}{r} dx|_{t=0} = -\int_{\mathbb{R}^{N}} \frac{(1+r)(n_{0}+\rho_{0})V_{0}}{r^{2}e^{r}} dx > 0.$$
(3.79)

Therefore,

$$H'(t) = H'(0) + \int_0^t \frac{d^2}{ds^2} H(s) ds \ge H'(0) > 0.$$
(3.80)

Multiplying the both sides of (3.78) by 2H'(t), we obtain

$$\left(H'(t)^{2}\right)' \geq \frac{1}{2^{\gamma-1}(1+\gamma)} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}}\right]^{\gamma-1} \left(H^{\gamma+1}(t)\right)'.$$
(3.81)

Integrate over [0, t], we obtain

$$H'(t)^{2} \geq \frac{1}{2^{\gamma-1}(1+\gamma)} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}} \right]^{\gamma-1} H^{\gamma+1}(t) + \left(\frac{d}{dt} \int_{\mathbb{R}^{N}} \frac{(n+\rho)}{re^{r}} dx \right)^{2}|_{t=0} - \frac{1}{2^{\gamma-1}(1+\gamma)} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}} \right]^{\gamma-1} H^{\gamma+1}(0).$$
(3.82)

Applying the assumption (3.7), we obtain

$$H'(t)^{2} \geq C_{0}^{2}H^{\gamma+1}(t) + \left(\frac{d}{dt}\int_{\mathbb{R}^{N}}\frac{(n+\rho)}{re^{r}}dx\right)^{2}|_{t=0} - C_{0}^{2}\left(\int_{\mathbb{R}^{N}}\frac{(n_{0}+\rho_{0})}{re^{r}}dx\right)^{\gamma+1} \geq C_{0}^{2}H^{\gamma+1}(t),$$
(3.83)

where $C_0 = \sqrt{\frac{1}{2^{\gamma-1}(1+\gamma)} \left[\frac{\Gamma(1+\frac{N}{2})}{(N-2)!\pi^{\frac{N}{2}}}\right]^{\gamma-1}}$. Therefore,

$$H'(t) \ge C_0 H^{\frac{\gamma+1}{2}}(t).$$
 (3.84)

By integrating over [0, t], we obtain

$$\int_{0}^{t} \frac{dH(s)}{H^{\frac{\gamma+1}{2}}(s)} \ge C_0 t.$$
(3.85)

Therefore,

$$-\frac{2}{\gamma-1}\left(H^{-\frac{\gamma-1}{2}}(t) - H^{-\frac{\gamma-1}{2}}(0)\right) \ge C_0 t, \tag{3.86}$$

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which means that

$$H(t) \ge \left(H^{-\frac{\gamma-1}{2}}(0) - \frac{\gamma-1}{2}C_0t\right)^{-\frac{\gamma}{\gamma-1}}.$$
(3.87)

From (3.6), we have

$$\int_{\mathbb{R}^N} \frac{n_0 + \rho_0}{re^r} dx > 0.$$
(3.88)

Applying the mass conservation, we obtain

$$H(t) = \int_{\mathbb{R}^{N}} \frac{n+\rho}{re^{r}} dx \le \int_{\mathbb{B}_{r_{0}}} \frac{n+\rho}{re^{r}} dx + \frac{1}{r_{0}} \int_{\mathbb{R}^{N}} (n_{0}+\rho_{0}) dx,$$
(3.89)

where \mathbb{B}_{r_0} is the N-dimensional ball centered at the origin with any given radius r_0 .

When $t \to \frac{2H(0)^{-\frac{\gamma-1}{2}}}{(\gamma-1)C_0}$, H(t) cannot be bounded. Therefore, we obtain that $\int_{r \le r_0} (n+\rho)(t,r)rdr$ cannot be bounded as $t \to \frac{2H(0)^{-\frac{\gamma-1}{2}}}{(\gamma-1)C_0}$, hence the solutions will blow up on or before the finite time $t \to \frac{2H(0)^{-\frac{\gamma-1}{2}}}{(\gamma-1)C_0}$. The proof is finished.

In the following, we will prove Theorem 3.3.

Proof of Theorem 3.3. We consider the solution (n, ρ, \vec{u}) satisfies the conditions in Theorem 3.3. From (3.67), we obtain

$$n_{tt} + \rho_{tt} = \Delta P + \nabla \cdot \left[\nabla \cdot \left((n+\rho)\vec{u} \otimes \vec{u}\right)\right].$$
(3.90)

Then, multiplying Eq (3.90) by $K_0(r)$ and taking the integration over \mathbb{R}^2 to obtain

$$\frac{d^2}{dt^2} \int_{\mathbb{R}^2} (n+\rho) K_0(r) dx = \int_{\mathbb{R}^2} \Delta P K_0(r) dx + \int_{\mathbb{R}^2} \nabla \cdot \left[\nabla \cdot \left((n+\rho) \vec{u} \otimes \vec{u} \right) \right] K_0(r) dx.$$
(3.91)

Using the same method as that we obtain from (3.69) to (3.73), we obtain from Lemma 3.5 that

$$\int_{\mathbb{R}^{2}} \Delta P K_{0}(r) dx = \int_{\mathbb{R}^{2}} (n^{\gamma} + \rho^{\gamma}) K_{0}(r) dx$$

$$\geq \frac{1}{2^{\gamma - 1} \left(\int_{\mathbb{R}^{2}} K_{0}(r) dx \right)^{\gamma - 1}} \left(\int_{\mathbb{R}^{2}} (n + \rho) K_{0}(r) dx \right)^{\gamma}.$$
(3.92)

Using the integration by parts, we obtain

$$\int_{\mathbb{R}^2} \nabla \cdot \left[\nabla \cdot \left((n+\rho) \vec{u} \otimes \vec{u} \right) \right] K_0(r) dx = \int_{\mathbb{R}^2} \sum_{i,j=1}^2 \frac{\partial^2 ((n+\rho)u_i u_j)}{\partial x_i \partial x_j} K_0(r) dx$$
$$= \int_{\mathbb{R}^2} (n+\rho) V^2 K_0''(r) dx > 0.$$
(3.93)

Therefore,

$$\frac{d^2}{dt^2}G(t) \ge \frac{1}{2^{\gamma-1} \left(\int_{\mathbb{R}^2} K_0(r) dx\right)^{\gamma-1}} G^{\gamma}(t), \tag{3.94}$$

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where $G(t) = \int_{\mathbb{R}^2} (n + \rho)(t, r) K_0(r) dx$. Using integration by parts, $(1.1)_1$, $(1.1)_2$ and (3.11), we obtain

$$\begin{aligned} G'(0) &= \frac{d}{dt} \int_{\mathbb{R}^2} (n+\rho) K_0(r) dx|_{t=0} \\ &= -\int_{\mathbb{R}^2} \nabla \cdot [(n+\rho) \vec{u}] K_0(r) dx|_{t=0} = \int_{\mathbb{R}^2} \sum_{i=1}^2 (n+\rho) u_i K_0'(r) \frac{x_i}{r} dx|_{t=0} \\ &= \int_{\mathbb{R}^2} \sum_{i=1}^2 (n+\rho) \frac{x_i}{r} V K_0'(r) \frac{x_i}{r} dx|_{t=0} = \int_{\mathbb{R}^2} (n+\rho) V K_0'(r) dx|_{t=0} \\ &= \int_{\mathbb{R}^2} (n_0+\rho_0) V_0 K_0'(r) dx > 0. \end{aligned}$$
(3.95)

Similarly, we obtain $G'(t) \ge 0$.

Multiplying both sides of (3.94) by 2G'(t), we obtain

$$\left(G'(t)^{2}\right)' \geq \frac{1}{2^{\gamma-1}(1+\gamma)\left(\int_{\mathbb{R}^{2}} K_{0}(r)dx\right)^{\gamma-1}} \left(G^{\gamma+1}(t)\right)'.$$
(3.96)

Integrate over [0, t] and using the assumption (3.11), we obtain

•

$$G'(t)^{2} \geq C_{1}^{2}G^{\gamma+1}(t) + \left(\frac{d}{dt}\int_{\mathbb{R}^{2}}(n+\rho)K_{0}(r)dx\right)^{2}|_{t=0} - C_{1}^{2}G^{\gamma+1}(0)$$

$$\geq C_{1}^{2}G^{\gamma+1}(t), \qquad (3.97)$$

where $C_1 = \sqrt{\frac{1}{2^{\gamma-1}(1+\gamma)}} \left(\int_{\mathbb{R}^2} K_0(r) dx \right)^{-\frac{\gamma-1}{2}}$. Therefore,

$$G'(t) \ge C_1 G^{\frac{\gamma+1}{2}}(t).$$
 (3.98)

By integrating over [0, t], we obtain

$$\int_{0}^{t} \frac{dG(s)}{G^{\frac{1+\gamma}{2}}(s)} \ge C_{1}t.$$
(3.99)

Therefore,

$$-\frac{2}{\gamma-1}\left(G^{-\frac{\gamma-1}{2}}(t) - G^{-\frac{\gamma-1}{2}}(0)\right) \ge C_1 t, \qquad (3.100)$$

which means that

$$G(t) \ge \left(G^{-\frac{\gamma-1}{2}}(0) - \frac{\gamma-1}{2}C_1 t\right)^{-\frac{\gamma}{\gamma-1}}.$$
(3.101)

From (3.10), we have

$$\int_{\mathbb{R}^2} (n_0 + \rho_0) K_0(r) dx > 0.$$
(3.102)

Applying the mass conservation, we obtain

$$G(t) = \int_{\mathbb{R}^2} (n+\rho) K_0(r) dx \le \int_{\mathbb{B}_{r_0}} (n+\rho) K_0(r) dx + \max_{r \ge r_0} K_0(r) \int_{\mathbb{R}^2} (n_0 + \rho_0) dx,$$
(3.103)

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where \mathbb{B}_{r_0} is the 2-dimensional ball centered at the origin with any given radius r_0 .

When $t \to \frac{2G(0)^{-\frac{\gamma-1}{2}}}{(\gamma-1)C_1}$, G(t) cannot be bounded. Thus, the solutions will blow up on or before the finite time $t \to \frac{2G(0)^{-\frac{\gamma-1}{2}}}{(\gamma-1)C_1}$. Therefore, we conclude that the solutions of the two-phase model in \mathbb{R}^2 will blow up.

4. Conclusions

In this paper, we study the blowup results of solutions to the two-phase model in \mathbb{R}^N . Our method depends on the energy integration method and a quasi-linear wave equation about ρ , the singularity of the two-phase model will form in finite time.

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Conflict of interest

On behalf of all authors, the corresponding author states that there is no conflict of interest.

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